Higgs theory, radiative corrections and current developments

Johannes Braathen (DESY)

7th International Workshop on "Higgs as a Probe of New Physics 2025" University of Osaka, Japan | 11 June 2025



QUANTUM UNIVERSE

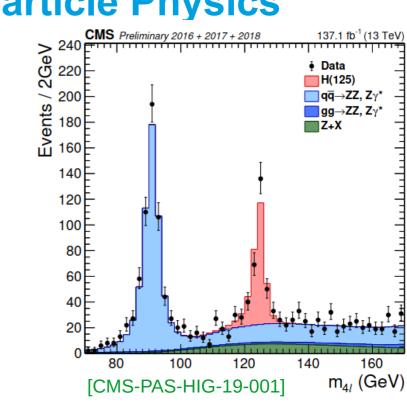






Higgs discovery in 2012: a milestone for Particle Physics

➤ 4th July 2012: discovery of <u>a Higgs boson</u> of mass 125 GeV by ATLAS and CMS collaborations at CERN Large Hadron Collider was a major milestone for Particle Physics

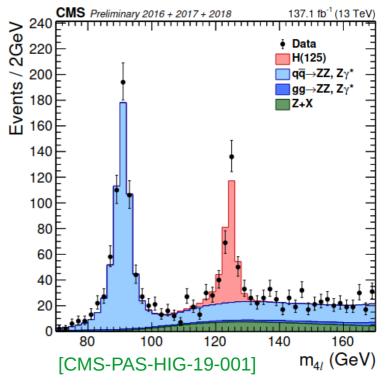


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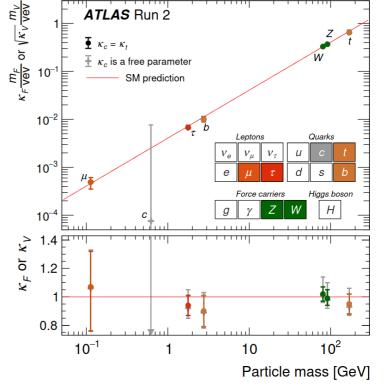
- Its mass M_h=125 GeV, to astonishing ~0.1% precision!
- The electroweak (EW) vacuum expectation value (vev) v=246 GeV
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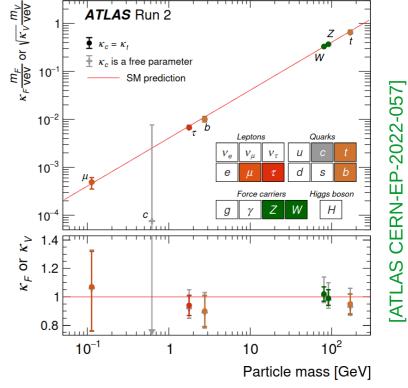
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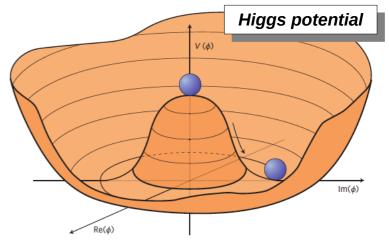
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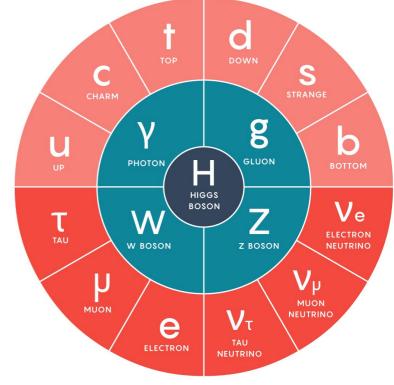


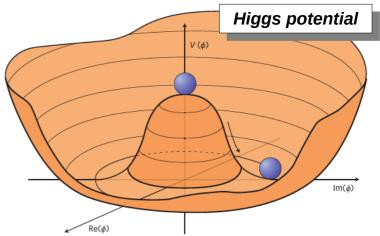
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- ➤ Particle content of **Standard Model** is "complete"
 - \rightarrow is this the end of the story?





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Open questions and motivation for New Physics

- What we still don't know about the Higgs boson:
 - Its couplings to 1st and 2nd gen. fermions
 - Its total width; BR(h→inv.) < 9%</p>
 - Its CP nature
 - Its fundamental nature (elementary or composite)
 - Structure of the Higgs sector (minimal or extended)
 - Form and origin of the Higgs potential (i.e. <u>why</u> do particles get masses, not just *how*)

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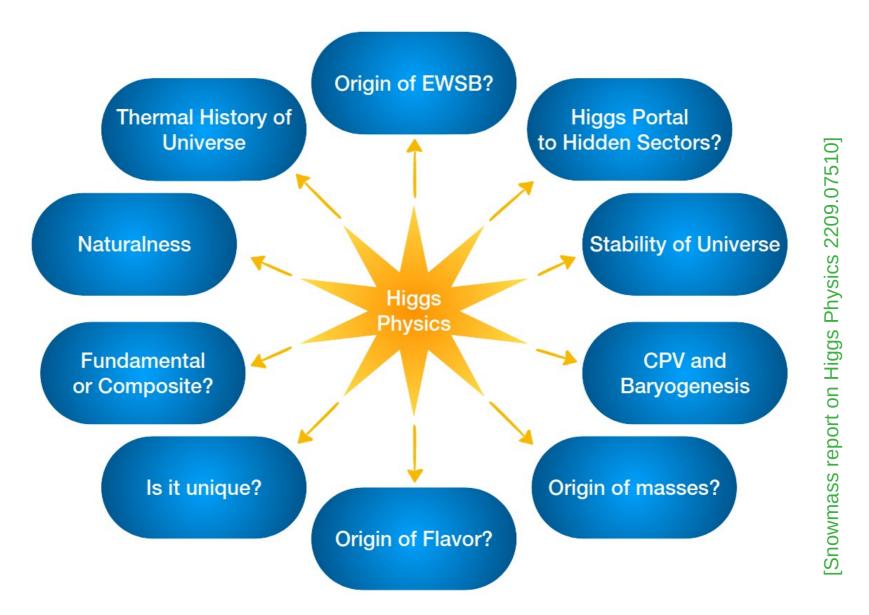
- Many further questions remain unanswered, e.g.
 - Gauge hierarchy problem, i.e. why is gravity so much weaker than the other forces (or why is the Planck scale so much higher than the electroweak scale)
 - Reason for 3 fermion families and origin of flavour
 - Origin of matter-antimatter asymmetry of the Universe
 - Dark Matter Etc.

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- ➤ Not addressed by our current best description of Particle Physics, the Standard Model (SM)
 - → New Physics must exist beyond-the-Standard-Model (BSM)!
- Many open problems relate to Higgs sector
 - → the 125-GeV Higgs boson will certainly play a key role in understanding the nature of BSM Physics
 - → BSM models often feature additional Higgs bosons/scalars

The Higgs boson plays a central role to probe New Physics



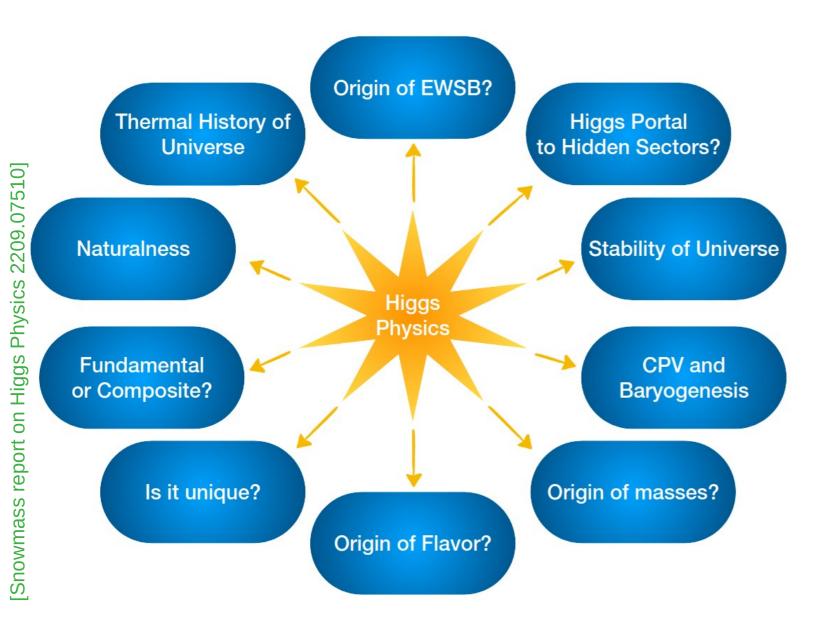
Outline of the lectures

- Part 1: Higgs as a Probe of New Physics
- Part 2: Some basics on radiative corrections and theory uncertainties
- Part 3: Higgs measurements and precision calculations
- Part 4: Di-Higgs production theory calculations, uncertainties and current developments
- Part 5: Future prospects (brief selection)

Probing New Physics with the Higgs boson

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The Higgs boson plays a central role to probe New Physics



Overview of this part:

- hierarchy problems
- form of the Higgs potential
- baryogenesis and electroweak phase transition

In backup:

- more details on hierarchy problems and their solutions
- Higgs portal to dark sectors
- Higgs as inflaton
- neutrino mass models with extended Higgs sectors

Hierarchy problems in Higgs Physics

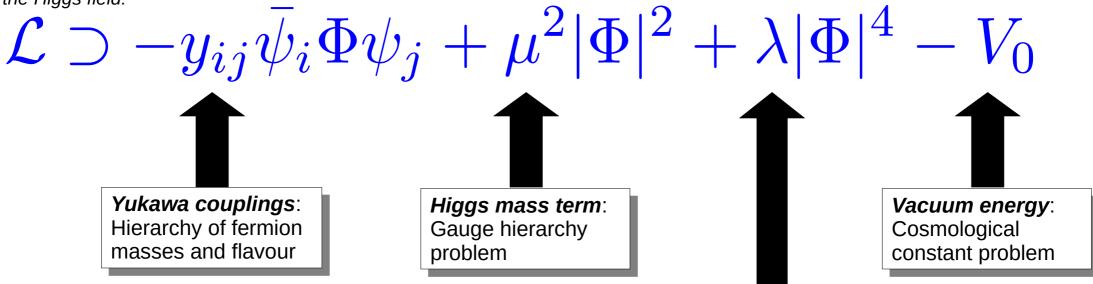
Parts of the SM Lagrangian involving only gauge bosons and fermions:

$$\mathcal{L} \supset -\frac{1}{4} F^a_{\mu\nu} F^{a,\mu\nu} + \bar{\psi}_i \gamma_\mu D^\mu_{ij} \psi_j$$

Slide adapted from [Salam '23], itself adapted from [Giudice]

→ entirely constrained by gauge symmetry, tested to high precision (e.g. at LEP)

Parts of the SM Lagrangian involving the Higgs field:

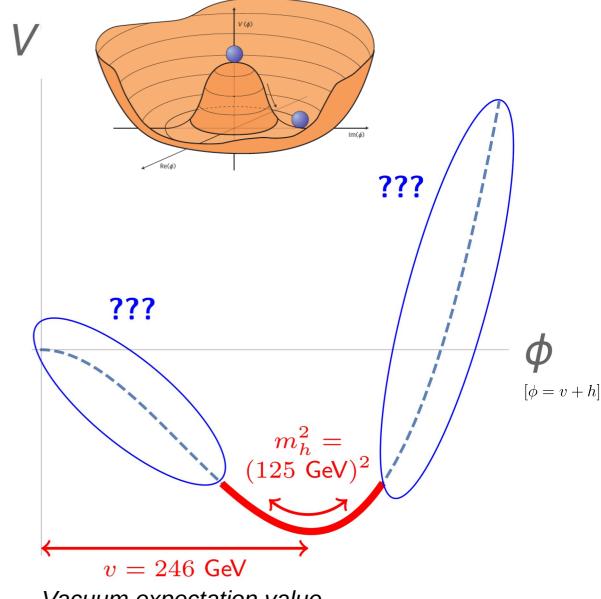


[More details on hierarchy problems, and on possible solutions to gauge hierarchy problem in backup]

Quartic Higgs coupling: UV behaviour and vacuum stability

Form of the Higgs potential and trilinear Higgs coupling

Brout-Englert-Higgs mechanism = origin of electroweak symmetry breaking ...
 ... but very little known about the Higgs potential causing the phase transition



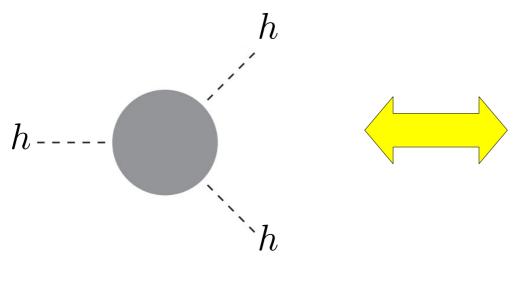
Vacuum expectation value

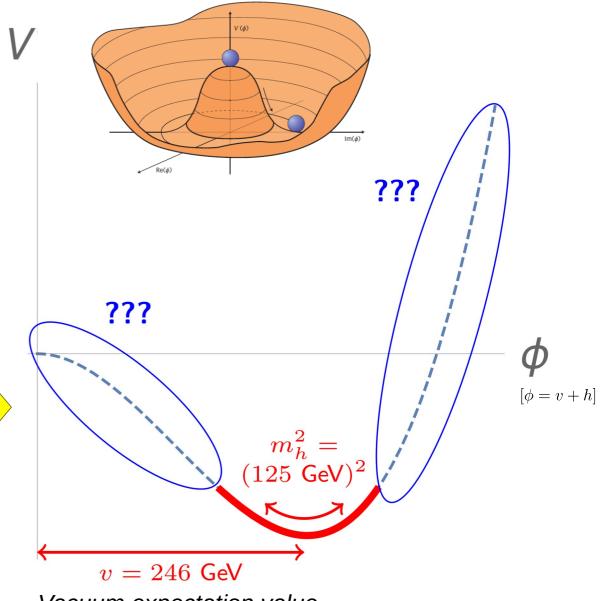
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> Trilinear Higgs coupling λ_{hhh} crucial to understand the shape of the potential





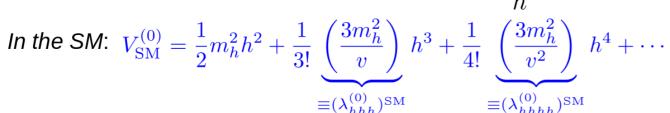
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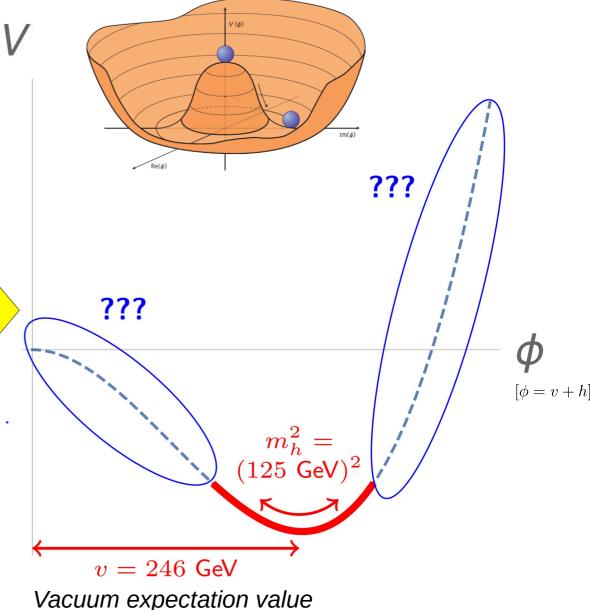
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Trilinear Higgs coupling λ_{hhh} crucial to understand the shape of the potential h



In general:

with
$$\kappa_{\lambda} \equiv \lambda_{hhh}/(\lambda_{hhh}^{(0)})^{\mathrm{SM}}$$
 and $\kappa_{4} \equiv \lambda_{hhhh}/(\lambda_{hhhh}^{(0)})^{\mathrm{SM}}$



Baryogenesis

Observed Baryon Asymmetry of the Universe (BAU)

$$\eta \equiv \frac{n_b - n_{\overline{b}}}{n_\gamma} \simeq 6.1 \times 10^{-10}$$
 [Planck '18]

 n_b : baryon no. density $n_{\overline{b}}$: antibaryon no. density n_v : photon no. density

- Sakharov conditions [Sakharov '67] for a theory to explain BAU:
 - 1) Baryon number violation
 - 2) C and CP violation
 - 3) Loss of thermal equilibrium

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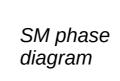
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- > **Sakharov conditions** [Sakharov '67] for a theory to explain BAU:
 - 1) Baryon number violation
 - 2) C and CP violation
 - 3) Loss of thermal equilibrium

→ Sphaleron transitions (break B+L)

→ C violation (SM is chiral), but not enough CP violation

 \rightarrow No loss of th. eq. \rightarrow in SM, the EWPT is a crossover



sym. phase

lst order

broken phase

2nd order crossover

m_H

75 GeV

SM cannot reproduce the BAU → BSM physics needed!

Electroweak Baryogenesis

- Many scenarios proposed, including:
 - Grand Unified Theories
 - Leptogenesis
 - Electroweak Baryogenesis (EWBG) [Kuzmin, Rubakov, Shaposhnikov, '85], [Cohen, Kaplan, Nelson '93]
- Sakharov conditions in EWBG
 - 1) Baryon number violation
- → Sphaleron transitions (break B+L)

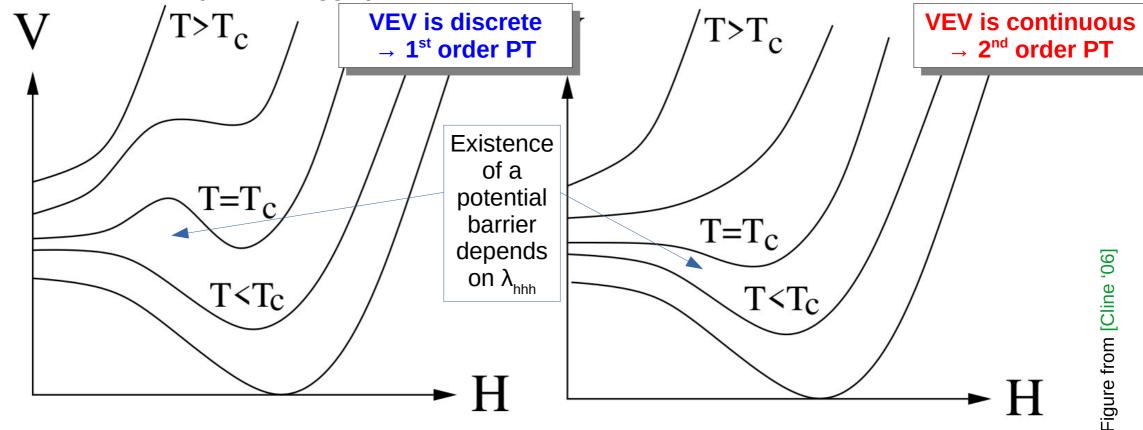
2) C and CP violation

- → C violation + CP violation in extended Higgs sector
- 3) Loss of thermal equilibrium
- → Loss of th. eq. via a strong 1st order EWPT

[More details on electroweak baryogenesis in backup]

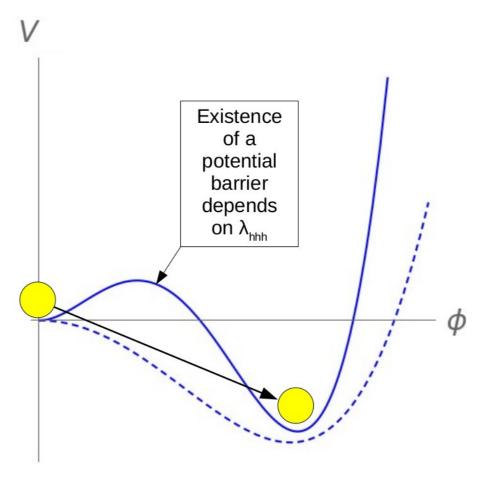
The Higgs potential and the Electroweak Phase Transition

Possible thermal history of the Higgs potential:



- \rightarrow λ_{hhh} determines the nature of the EWPT!
 - \Rightarrow deviation of λ_{hhh} from its SM prediction typically* needed to have a strongly first-order EWPT [Grojean, Servant, Wells '04], [Kanemura, Okada, Senaha '04]

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Radiative corrections 101

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Impact of radiative corrections: W-boson mass

- Electroweak precision observables, such as
 - W-boson mass M_w
 - > Effective leptonic weak mixing angle $\rightarrow \sin^2 \theta_{\text{eff}}^{\text{lep}}$
 - Z-boson decay width Γ₇
 - Muon anomalous magnetic moment (g-2)_µ
 etc.

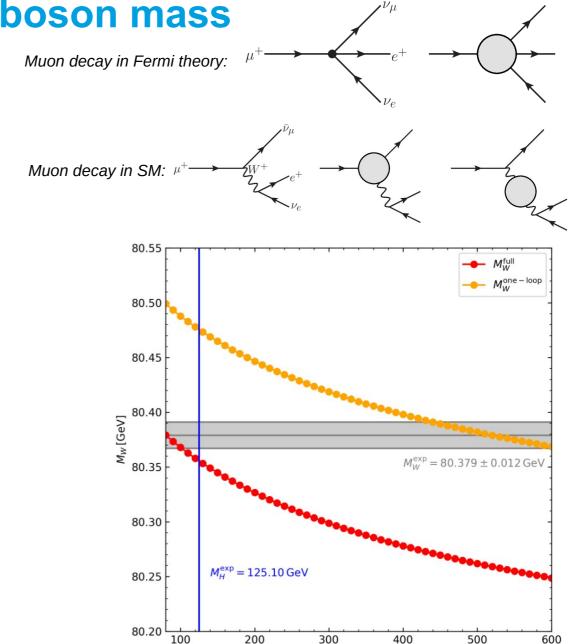
are measured very precisely, and can also be computed to high level of accuracy in terms of G_F , $\alpha(0)$, M_Z (most precisely measured EW quantities) and M_h , M_t , α_S , $\Delta\alpha^{had}$, $\Delta\alpha^{lept}$, m_h , etc.

 \triangleright Relation between M_W, M_Z, G_F, $\alpha(0)$ obtained by matching the calculation of muon decay between the SM and Fermi theory

$$\frac{G_F}{\sqrt{2}} = \frac{e^2}{8M_W^2 s_W^2} (1 + \Delta r) \quad \Rightarrow \quad M_W^2 \left(1 - \frac{M_W^2}{M_Z^2} \right) = \frac{\pi \alpha}{\sqrt{2} G_F} (1 + \Delta r)$$

with $\Delta r \equiv \Delta r(M_W, M_Z, M_h, M_t, ...)$ the corrections to muon decay (w/o finite QED effects)

- Without Δr (loop corrections), M_w ~ 80.9 GeV, i.e. ~40σ away from experimental measurement! One-loop calculation also ~10σ off \rightarrow incorporation of (known) higher orders is essential
- Allows testing the SM as well as BSM models at quantum level DESY. | Lecture on Higgs theory, HPNP | Johannes Braathen (DESY) | 11 June 2025



 M_{μ}^{SM} [GeV]

Figure by G. Weiglein

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Impact of radiative corrections: h→bb decay

ightharpoonup Decay h → bb already at tree level (driven by bottom Yukawa coupling, prop. to m_b)

$$\Gamma^{(0)}(h \to b\bar{b}) \stackrel{M_h \gg m_b}{=} \frac{3G_F}{4\sqrt{2}\pi} M_h m_b^2$$
 (1)

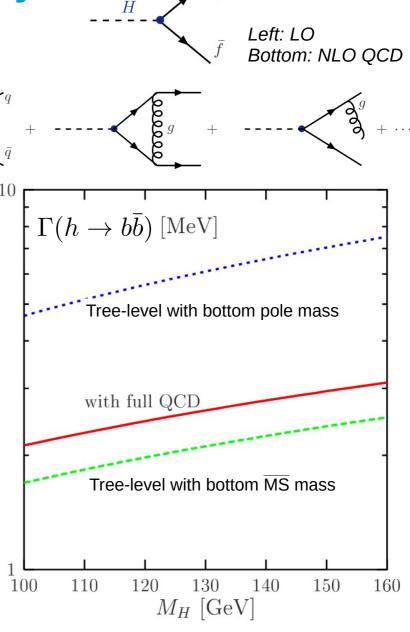
 \triangleright Large QCD corrections (driven by strong gauge coupling α_s which does not enter at tree level)

$$\hat{\Gamma}(h \to b\bar{b}) \stackrel{M_h \gg m_b}{=} \frac{3G_F}{4\sqrt{2}\pi} M_h m_b^2 \left[1 + \frac{4}{3} \frac{\alpha_s}{\pi} \left(\frac{9}{4} + \frac{3}{2} \log \frac{m_b^2}{M_h^2} \right) + \cdots \right]$$
 (2)

also contains large logs $log(m_b/M_h) \rightarrow can \ spoil \ perturbative$ expansion

> Can be resummed to all orders in α_s (\rightarrow c.f. Effective Field Theories) by expressing the decay width in terms of $\overline{\text{MS}}$ bottom mass

$$\hat{\Gamma}(h \to b\bar{b}) \stackrel{M_h \gg m_b}{=} \frac{3G_F}{4\sqrt{2}\pi} M_h \left(\bar{m}_b^{\overline{\text{MS}}}(M_h)\right)^2 \left[1 + 5.67 \frac{\alpha_s(M_h)}{\pi} + \cdots\right]$$
 (3)



Theoretical uncertainties: sources and estimates

 \triangleright Meaningful theory calculations relate physical observables (e.g. M_W , M_Z , α_{em} , G_F , ...)

Let's write in general: $O = \mathcal{F}(I_1, I_2, \cdots)$

with I_1 , I_2 , ... input parameters

- > 2 main sources of theoretical uncertainties:
 - (1) Unknown higher-order and/or subleading contributions (in some step of the calculation)
 - Example 1: O is computed with a fixed-order calculation as

$$O = \underbrace{\mathcal{F}^{(0)}(I_1,I_2,\cdots)}_{\text{tree level}} + \underbrace{\frac{1}{16\pi^2}\mathcal{F}^{(1)}(I_1,I_2,\cdots)}_{\text{one loop}} + \underbrace{\frac{1}{(16\pi^2)^2}\mathcal{F}^{(2)}(I_1,I_2,\cdots)}_{\text{two loops}} + \cdots$$

- → perturbative expansion *truncated* at some order, higher-orders are unknown
- <u>Example 2</u>: O computed with an Effective-Field-Theory calculation, in terms of parameter g_A obtained from EFT matching, i.e. $O = \tilde{\mathcal{F}}_{EFT}^{(0)}(I_1,g_A,\cdots) + \frac{1}{16\pi^2}\tilde{\mathcal{F}}_{EFT}^{(1)}(I_1,g_A,\cdots) + \cdots$

with EFT matching of g_A : $g_A = \mathcal{G}^{(0)}(I_1, I_2, \cdots, \Lambda_{\text{match.}}) + \frac{1}{16\pi^2} \mathcal{G}^{(1)}(I_1, I_2, \cdots, \Lambda_{\text{match.}})$

 \rightarrow missing higher-orders in EFT calc. + in matching + uncertainty from choice of matching scale Λ_{match}

How to estimate their effect? → renormalisation scheme conversions

→ variations of renormalisation scale or matching scale

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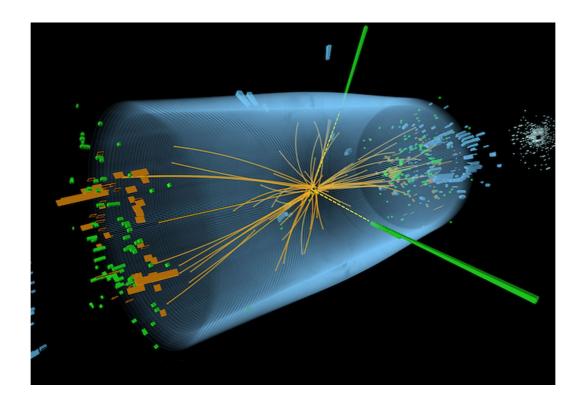
- (2) Finite precision with which input parameters are known
 - \rightarrow Assuming (for simplicity) that I₁, I₂, ... are physical observables, their values are obtained from experimental measurements, that have an error/uncertainty

How to estimate this effect? → error propagation (or simply repeat calculation with varied inputs)

e.g.
$$\Delta O = \sqrt{\left(\frac{\partial \mathcal{F}}{\partial I_1}\right)^2 (\Delta I_1)^2 + \left(\frac{\partial \mathcal{F}}{\partial I_2}\right)^2 (\Delta I_2)^2 + \cdots}$$

Experimental errors on I_1 , I_2 , ...

Higgs measurements and precision calculations

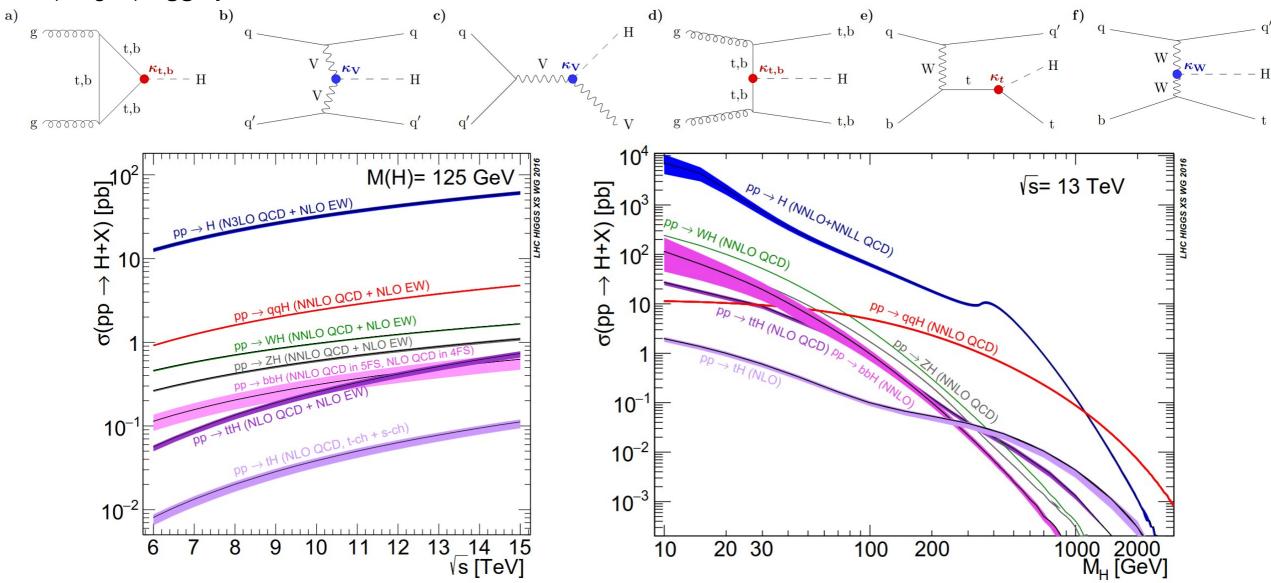


 $A h \rightarrow yy \text{ event at}$ CMS

Higgs production at LHC

Diagrams from [CMS Nature '22], Plots from [LHC Higgs WG '16] See also reviews of [Djouadi '05]

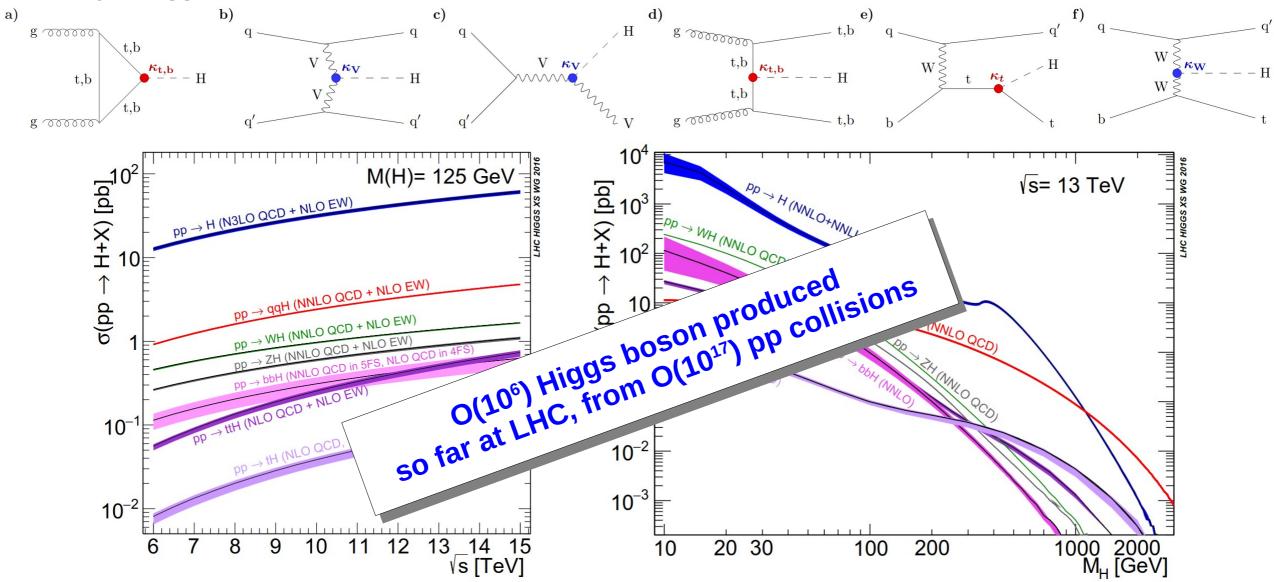
> (Single-)Higgs production channels at LHC



Higgs production at LHC

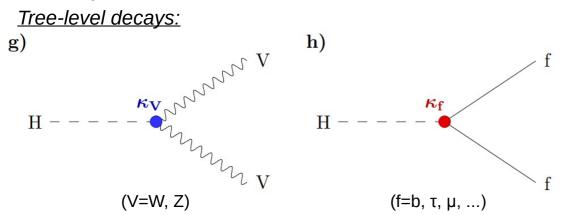
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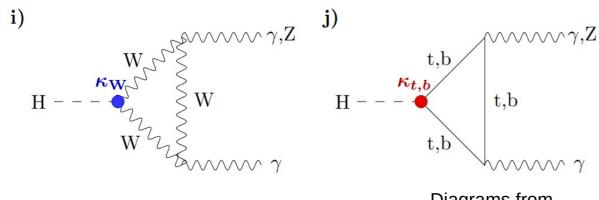


Higgs decay channels

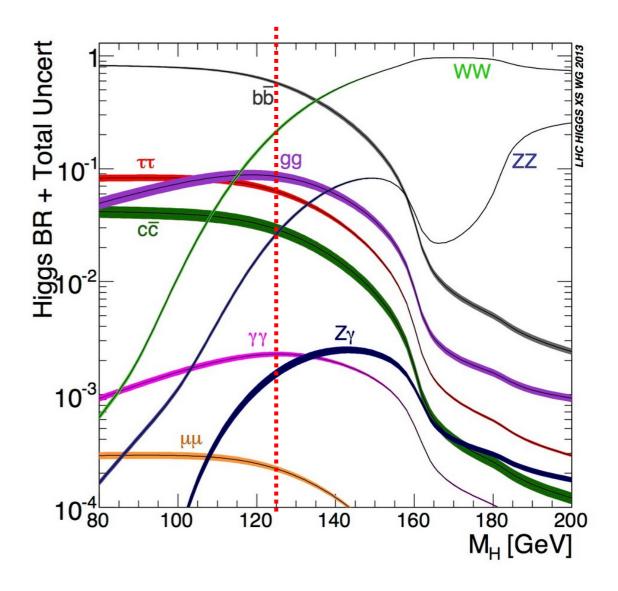
Decay channels



<u>Loop-induced decays:</u>

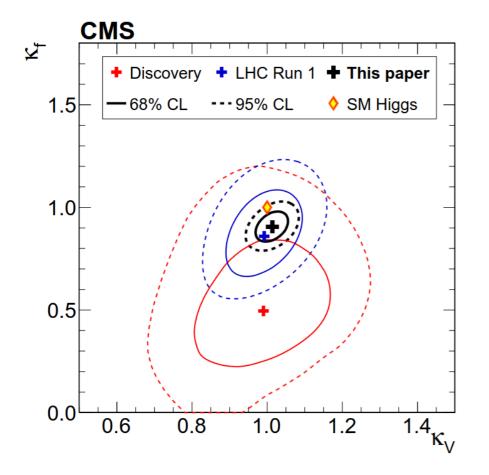


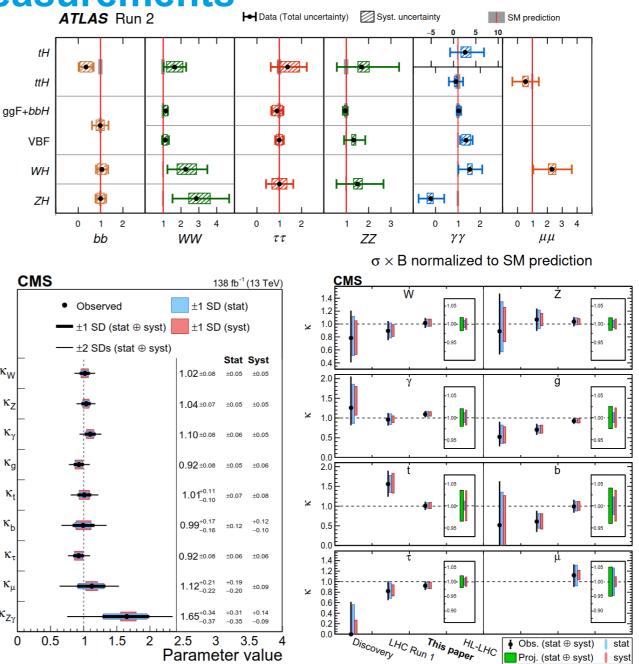
Diagrams from [CMS Nature '22]



Example results of Higgs measurements

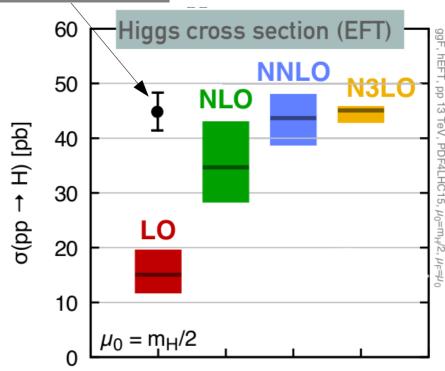
Example results from ATLAS and CMS 2022
"10-year of the Higgs discovery" Nature papers,
for coupling modifiers (bottom)
or signal strengths (right)





Interpreting experimental results

- Comparison between experiment and theory carried out at the level of:
 - Signal strengths
 - κ parameters (signal strength modifiers)
 - Simplified Template Cross-Sections (STXS)
 - Fiducial cross sections
 - Coefficients of EFT operators
- Requires **high-precision theoretical predictions** (with level of accuracy at least matching that of experimental results)
 - → both in SM and BSM theories
 - → huge efforts from precision calculation communities (QCD, EW, BSM)



Exp. measurement

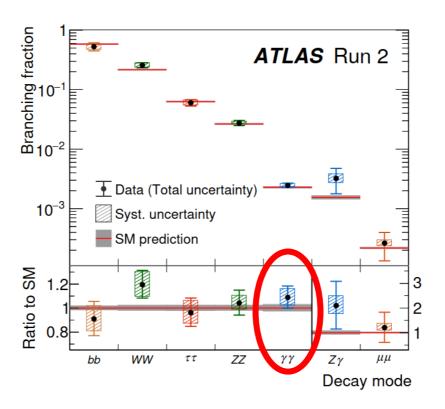
Total cross section for (inclusive) single-Higgs production, in heavy top limit $(m_t \rightarrow +\infty)$

Figure taken from [Weiglein '22], itself from [Wiesemann '22], based on results from [Anastasiou et al. '15], [Mistlberger '18]

Public tools to confront model predictions with experimental results:

- → HiggsSignals (signal strengths, STXS) [Bechtle et al '13, '20] → now included in HiggsTools [Bahl et al '22]
- Lilith (signal strengths) [Bernon, Dumont '15], [Kraml et al '19], [Bertrand et al '20]

An example calculation of Higgs properties in a BSM model: leading two-loop corrections to $\Gamma(h \rightarrow \gamma \gamma)$ in the Inert Doublet Model



DESY.

The Inert Doublet Model

> 2 SU(2), doublets Φ_{12} of hypercharge $\frac{1}{2}$

$$\Phi_1 = \begin{pmatrix} G^+ \\ \frac{1}{\sqrt{2}}(v+h+iG) \end{pmatrix} \quad \text{and} \quad \Phi_2 = \begin{pmatrix} H^+ \\ \frac{1}{\sqrt{2}}(H+iA) \end{pmatrix}$$

$$\Phi_2 = \begin{pmatrix} H^+ \\ \frac{1}{\sqrt{2}}(H + iA) \end{pmatrix}$$

► Unbroken Z, symmetry $\Phi_1 \rightarrow \Phi_1$, $\Phi_2 \rightarrow -\Phi_2$

$$V_{\text{IDM}}^{(0)} = \mu_1^2 |\Phi_1|^2 + \mu_2^2 |\Phi_2|^2 + \frac{\lambda_1}{2} |\Phi_1|^4 + \frac{\lambda_2}{2} |\Phi_2|^4 + \lambda_3 |\Phi_1|^2 |\Phi_2|^2 + \lambda_4 |\Phi_2^{\dagger} \Phi_1|^2 + \frac{\lambda_5}{2} \left((\Phi_2^{\dagger} \Phi_1)^2 + \text{h.c.} \right)$$

- Inert scalars H, A, H[±]: charged under Z₂ symmetry (Z₂-odd)
- Model parameters: 3 BSM masses m_{H} , m_{A} , m_{H+} , BSM mass scale μ_{2} , inert doublet quartic self-coupling λ_{2}
- Lightest inert scalar = Dark Matter candidate
 - → assume H here

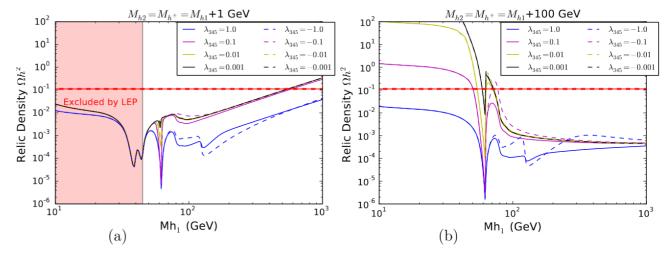
Drawing by [K. Radchenko

Dark Matter in the Inert Doublet Model

- DM (H) relic density obtained via freeze-out mechanism, while evading current detection bounds
- 2 possible scenarios:
 - → "Higgs resonance scenario" m_H~m_h/2
 - → "Heavy Higgs scenario" m_H≥500 GeV
- IDM testable at current and future experiments via
 - DM direct and indirect searches
 - direct searches at colliders

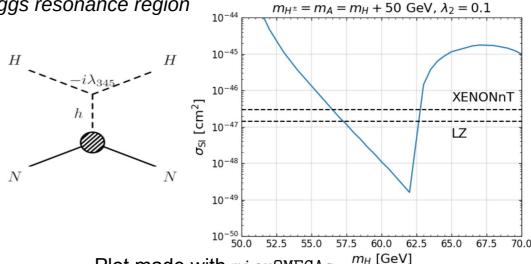
(see also [JB, Gabelmann, Robens, Stylianou '24])

- precision/indirect tests
 - → properties of h₁₂₅



[Belyaev et al. '16]

Direct detection bounds around Higgs resonance region



Plot made with micrOMEGAs

[Bélanger et al. '18]

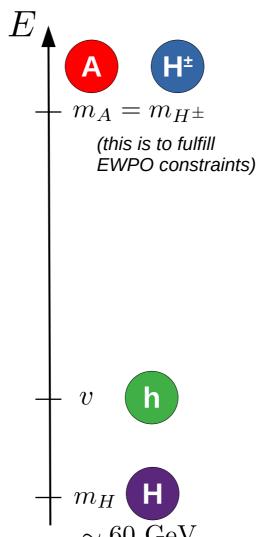
Dark Matter in the Inert Doublet Model

- DM (H) relic density obtained via freeze-out mechanism, while evading current detection bounds
- 2 possible scenarios:
 - → "Higgs resonance" scenario m_H~m_h/2
 - → "Heavy Higgs" scenario m_H≥500 GeV
- IDM testable at current and future experiments via
 - DM direct and indirect searches
 - direct searches at colliders

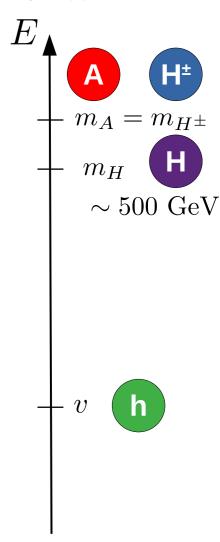
(see also [JB, Gabelmann, Robens, Stylianou '24])

- precision/indirect tests
 - → properties of h₁₂₅

"Higgs resonance" scenario



"Heavy Higgs" scenario



Higgs decay to two photons: existing one-loop results

- DM scenarios of IDM investigated via Higgs properties at one loop (1L) in [Kanemura, Kikuchi, Sakurai '16]
- Additional charged inert Higgs \rightarrow Higgs decay to 2 photons especially important!

$$\Gamma[h \to \gamma \gamma] \simeq \frac{\sqrt{2} G_F \alpha_{\rm EM}^2 m_h^3}{64\pi^3} \left| -\frac{1}{6} \left(1 - \frac{\mu_2^2}{m_{H^{\pm}}^2} \right) + \sum_f Q_f^2 N_c^f I_f[m_h^2] + I_W[m_h^2] \right|^2 \qquad \text{with } m_{H^{\pm}}^2 = \mu_2^2 + \lambda_3 v^2 \,,$$

<u></u>

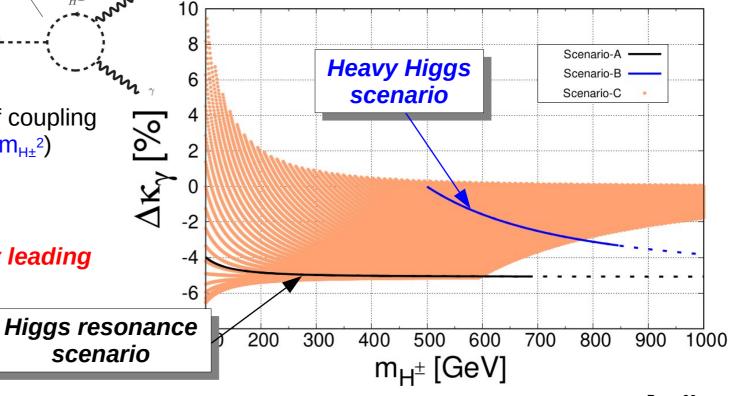
scenario

 I_f , I_W : fermion/W-boson loops (SM-like)

Charged Higgs contribution:

Compensation between mass dependence of coupling $(\lambda_3=2(m_{H\pm}^2-\mu_2^2)/v^2)$ and of loop function $(C_0\sim1/m_{H\pm}^2)$

- \rightarrow does not decouple when increasing m_{H±}!
- $^{>}$ h → γγ is a loop-induced decay, i.e. **1L** is only leading order (LO)
 - → What about 2L (NLO) corrections?



Higgs Low-Energy Theorem

- Calculation of 2L 3-point functions with external momenta not possible in general
- Assuming m_h << heavy BSM scalar masses, we can employ a Higgs Low-Energy Theorem (see e.g. [Kniehl, Spira '95])
- Compute effective Higgs-photon coupling C_{hyy} of the form

$$\mathcal{L}_{\text{eff}} = -\frac{1}{4} C_{h\gamma\gamma} h F^{\mu\nu} F_{\mu\nu}$$

by taking derivative of (unrenormalised) photon self-energy w.r.t Higgs field

$$C_{h\gamma\gamma} = \frac{\partial}{\partial h} \Pi_{\gamma\gamma} (p^2 = 0) \bigg|_{h=0} \quad \text{where } \Sigma_{\gamma\gamma}^{\mu\nu} (p^2) = (p^2 g^{\mu\nu} - p^{\mu} p^{\nu}) \Pi_{\gamma\gamma} (p^2)$$

- Schematically: $\frac{\partial}{\partial h} \left[- - - \right] = \frac{\int_{---}^{+--}^{+--} h(p^2 = 0)}{\int_{----}^{+---}^{+---} h(p^2 = 0)}$
- Neglects incoming momentum on Higgs leg, but valid for m_h << m_{H,A,H±}

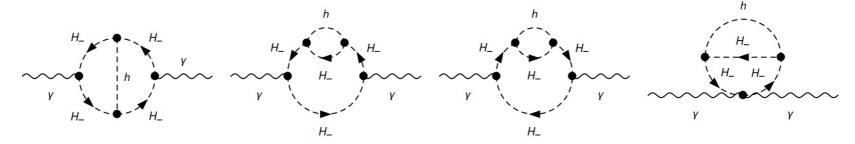
Computing two-loop BSM corrections to Γ(h → yy)

- All known SM contributions:
 - QCD up to 3L [Djouadi '08] (+ refs. therein)
 - EW SM-like to full 2L [Degrassi, Maltoni '05], [Actis et al. '09]
- Our new calculation: leading two-loop BSM contributions
 - genuine, dominant, 2L contributions involving inert scalars (+ SM-like scalars and/or gauge bosons)
 - purely scalar and fermion-scalar contributions to (1L)^2 terms from external-leg and VEV renormalisation

$$C_{h\gamma\gamma}^{(2), \text{ IDM}} = C_{h\gamma\gamma}^{\mathcal{O}(\lambda_3^2)} + C_{h\gamma\gamma}^{\mathcal{O}((\lambda_4 + \lambda_5)^2)} + C_{h\gamma\gamma}^{\mathcal{O}((\lambda_4 - \lambda_5)^2)} + C_{h\gamma\gamma}^{\mathcal{O}(\lambda_2)} + C_{h\gamma\gamma}^{\text{ext.-leg.+VEV}}$$

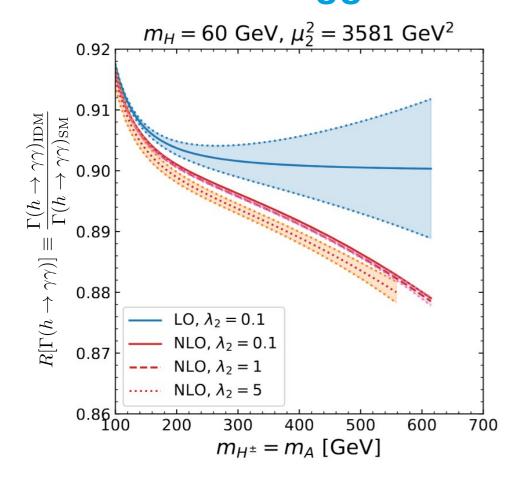
Example:

 $O(\lambda_3^2)$ diagrams

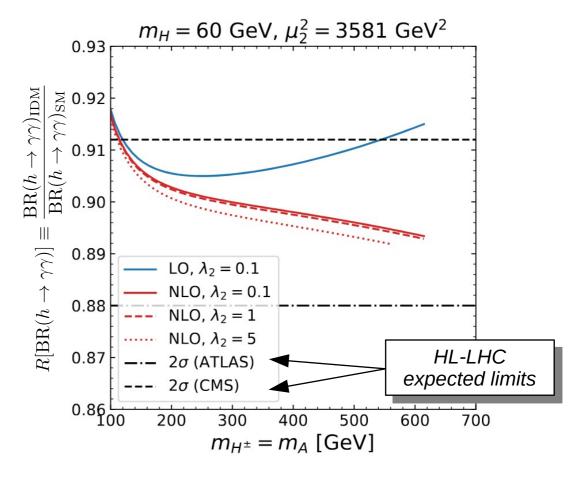


Photon self-energy diagrams generated with FeynArts, computed with FeynCalc and Tarcer, reduced to (limits of) integrals known analytically; then derivative w.r.t. h taken

Results for the Higgs resonance scenario



Inclusion of two-loop (NLO) corrections significantly reduces the theoretical uncertainty

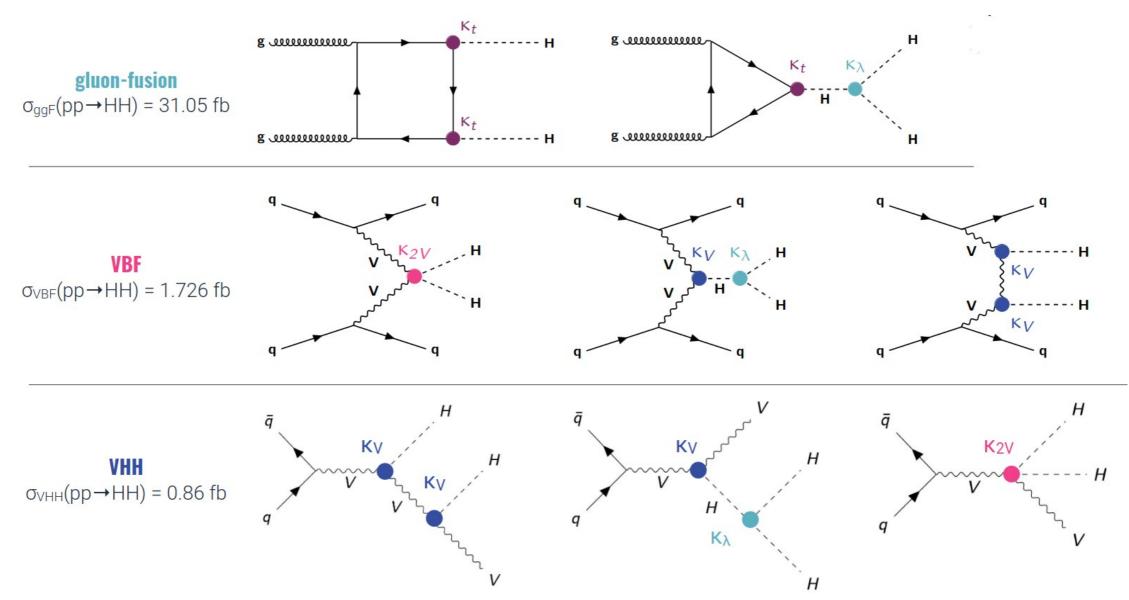


- **Almost entire scenario** (for $m_{H\pm}$ > 120 GeV) can be ruled out if no deviation is found in h → yy!
- Proper interpretation of experimental results requires inclusion of two-loop corrections!

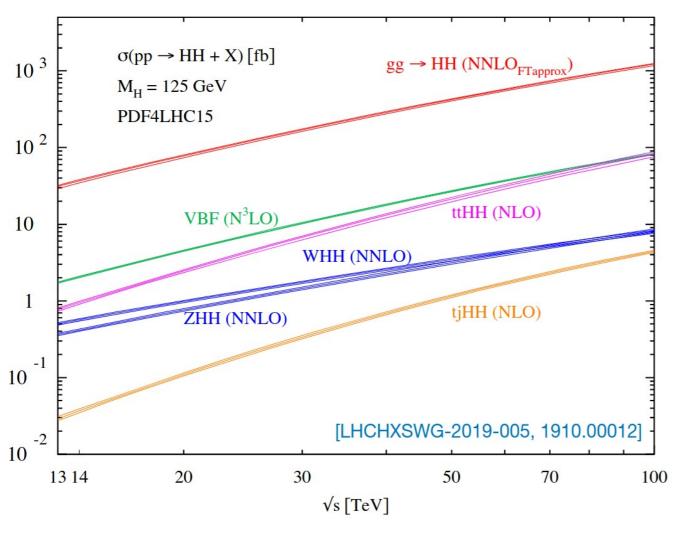
Di-Higgs production: Theory predictions and uncertainties

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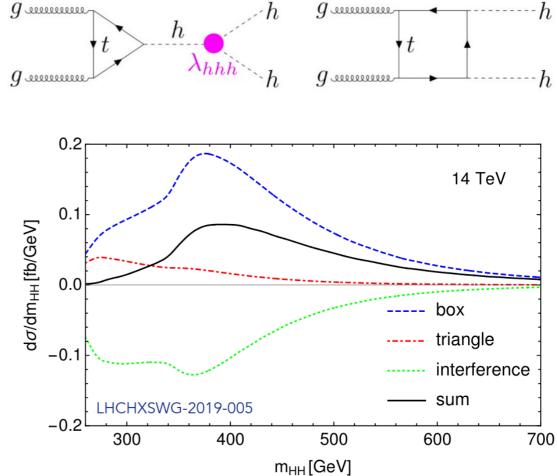
Different channels for di-Higgs production



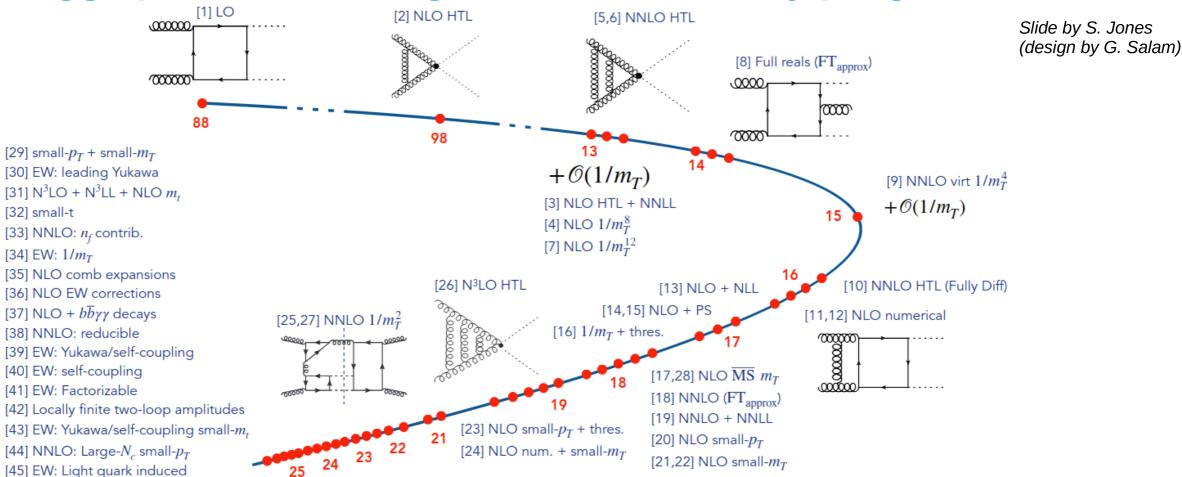
Different channels for di-Higgs production



Leading channel: gluon fusion



Di-Higgs production via gluon fusion: theory progress in SM



[1] Glover, van der Bij 88; [2] Dawson, Dittmaier, Spira 98; [3] Shao, Li, Li, Wang 13; [4] Grigo, Hoff, Melnikov, Steinhauser 13; [5] de Florian, Mazzitelli 13; [6] Grigo, Melnikov, Steinhauser 14; [7] Grigo, Hoff, Steinhauser 14; [7] Grigo, Hoff, Steinhauser 15; [10] de Florian, Grazzini, Hanga, Kallweit, Lindert, Maierhöfer, Mazzitelli, Rathlev 16; [11] Borowka, Greiner, Heinrich, Jones, Kerner, Schlenk, Schubert, Zirke 16; [12] Borowka, Greiner, Heinrich, Jones, Kerner, Schlenk, Schubert, Zirke 16; [12] Borowka, Greiner, Heinrich, Jones, Kerner, Schlenk, Schubert, Zirke 16; [13] Ferrera, Pires 16; [14] Heinrich, Jones, Kerner, Luisoni, Vryonidou 17; [15] Jones, Kuttimalai 17; [16] Gröber, Maier, Rauh 17; [17] Baglio, Campanario, Glaus, Mühlleitner, Spira, Streicher 18; [18] Grazzini, Heinrich, Jones, Kallweit, Kerner, Lindert, Mazzitelli 18; [19] de Florian, Mazzitelli 18; [20] Bonciani, Degrassi, Giardino, Gröber 18; [21] Davies, Mishima, Steinhauser, Wellmann 18, 18; [22] Mishima 18; [23] Gröber, Maier, Rauh 19; [24] Davies, Heinrich, Jones, Kerner, Mishima, Steinhauser, Davied Wellmann 19; [25] Davies, Steinhauser 19; [26] Chen, Li, Shao, Wang 19, 19; [27] Davies, Herren, Mishima, Steinhauser 19, 21; [28] Baglio, Campanario, Glaus, Mühlleitner, Ronca, Spira 21; [29] Bellafronte, Degrassi, Giardino, Gröber, Vitti 22; [30] Davies, Mishima, Schönwald, Steinhauser, Zhang 22; [31] Ajjath, Shao 22; [32] Davies, Mishima, Schönwald, Steinhauser, Zhang 23; [35] Bagnaschi, Degrassi, Gröber 23; [36] Bi, Huang, Huang, Ma Yu 23 [37] Li, Si, Wang, Zhang, Zhao 24; [43] Davies, Schönwald, Steinhauser, Zhang 25; [44] Davies, Schönwald, Steinhauser, Zhang 25; [45] Bonetti, Rendler, Bobadilla 25;

Di-Higgs production via gluon fusion: uncertainty budget

Combination of NLO and N^mLO HTL yields:

• Scale uncertainty of: +2.1% / -4.9%

• PDF + α_s : $\pm 2.2 \%$

• m_T approx: $\pm 2.7 \%$

HWG HH Twiki

[Chen, Li, Shao, Wang 19, 19; Grazzini, Heinrich, Jones, Kallweit, Kerner, Lindert, Mazzitelli 18; de Florian, Grazzini, Hanga, Kallweit, Lindert, Maierhöfer, Mazzitelli, Rathlev 16; Maltoni, Vryonidou, Zaro 14; Borowka, Greiner, Heinrich, Jones, Kerner, Schlenk, Schubert, Zirke 16; Dawson, Dittmaier, Spira 98; Glover, van der Bij 88]

Converting the top quark mass to the \overline{MS} scheme

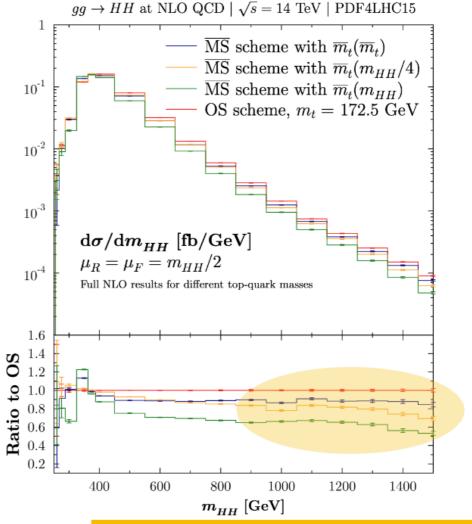
$$m_t \to \overline{m}_t(\mu) \left(1 + \frac{\alpha_s(\mu)}{4\pi} C_F \left\{ 4 + 3 \log \left[\frac{\mu^2}{\overline{m}_t(\mu)^2} \right] \right\} \right)$$

This leads to an additional uncertainty related to the choice of the top-quark mass scheme

$$\sqrt{s} = 13 \text{ TeV}: \quad \sigma_{tot} = 27.73(7)^{+4\%}_{-18\%} \text{ fb},$$

 $\sqrt{s} = 14 \text{ TeV}: \quad \sigma_{tot} = 32.81(7)^{+4\%}_{-18\%} \text{ fb},$

Currently: attempts at resumming these large logs by EFT for high-energy limit (SCET = soft collinear effective theory) → [Kaskiewicz et al. '25] → improvement but new uncertainties from matching scale

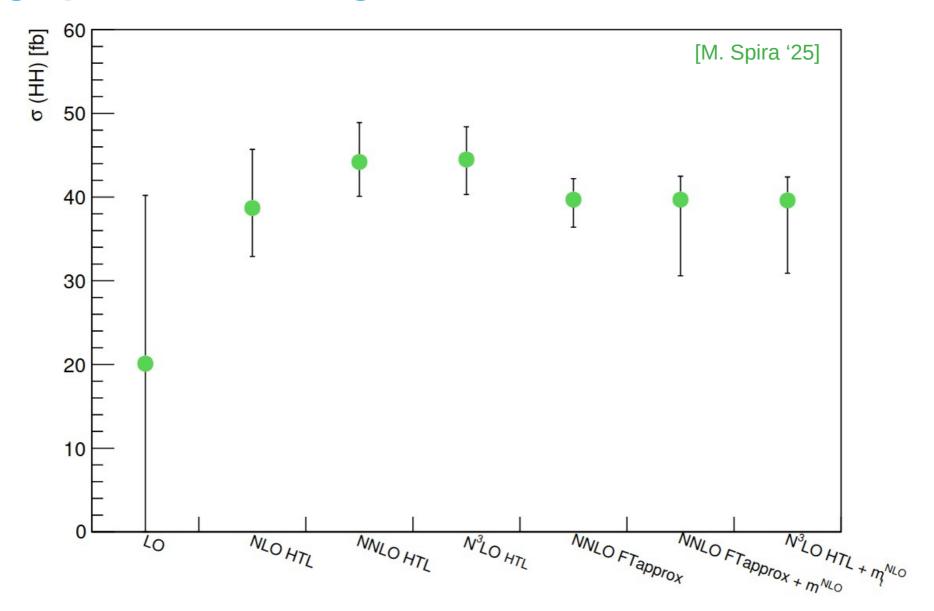


Large uncertainties in the high-energy limit

[J. Baglio, F. Campanario, S. Glaus, M. Muehlleitner, J. Ronca, M. Spira, J. Streicher, 2003.03227, 2008.11626]

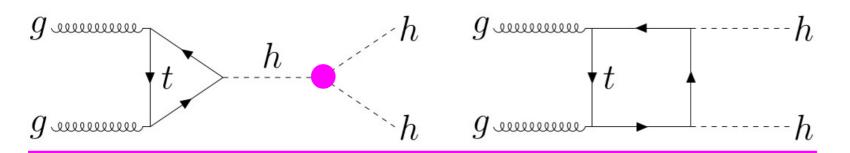
[Slide elements by S. Jaskiewicz]

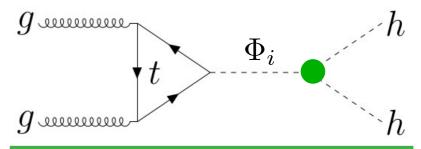
Di-Higgs production via gluon fusion: evolution of uncertainty



Di-Higgs production in BSM models

Leading order (LO) diagrams (involving top quark) in BSM models





"Non-resonant contributions"

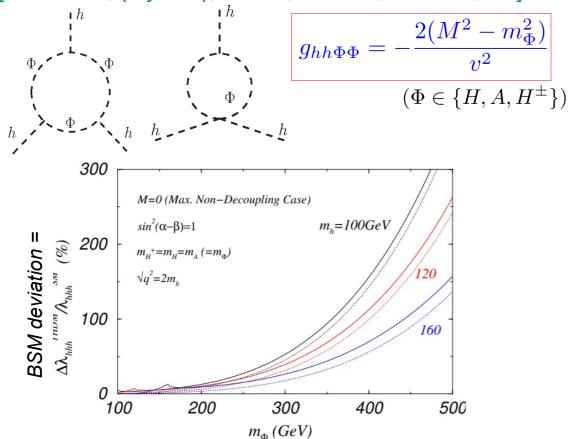
- Standard Model (SM)-like diagrams
- Involves the trilinear self-coupling of $h_{125} \lambda_{hhh}$
 - → probe of the Higgs potential
- Large destructive interference between triangle and box diagram
 - → suppression of cross-section in SM
 - ightarrow large changes in di-Higgs cross-section possible from BSM effects in λ_{hhh}

"Resonant contributions"

- Diagrams involving BSM scalars in s-channel (here generically denoted Φ_i)
- → collider searches for BSM scalars
- Involve BSM trilinear scalar couplings λ_{iik}
- → probe of Higgs potential in extended scalar sectors

Mass splitting effects in λ_{hhh}

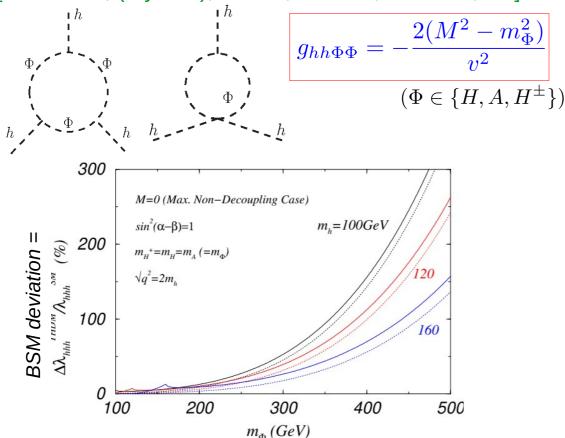
First investigation of 1L BSM contributions to λ_{hhh} in 2HDM: [Kanemura, (Kiyoura), Okada, Senaha, Yuan '02, '04]



- Deviations of tens/hundreds of % from SM possible, for large $g_{h\Phi\Phi}$ or $g_{hh\Phi\Phi}$ couplings
- Mass splitting effects, now found in various models (2HDM, inert doublet model, singlet extensions, etc.)

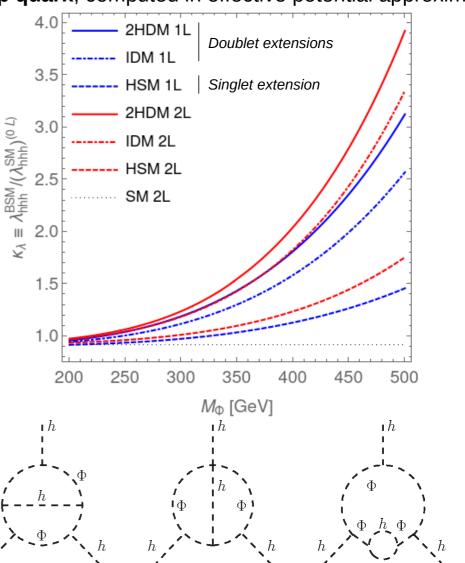
Mass splitting effects in λ_{hhh}

First investigation of 1L BSM contributions to λ_{hhh} in 2HDM: [Kanemura, (Kiyoura), Okada, Senaha, Yuan '02, '04]

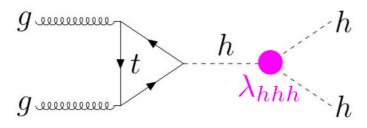


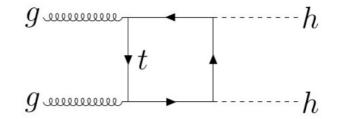
- **Deviations of tens/hundreds of % from SM possible,** for large $g_{h\Phi\Phi}$ or $g_{hh\Phi\Phi}$ couplings
- Mass splitting effects, now found in various models (2HDM, inert doublet model, singlet extensions, etc.)

- Large effects confirmed at 2L in [JB, Kanemura '19]
- → leading 2L corrections involving BSM scalars (H,A,H±) and top quark, computed in effective potential approximation



Interference in non-resonant di-Higgs production



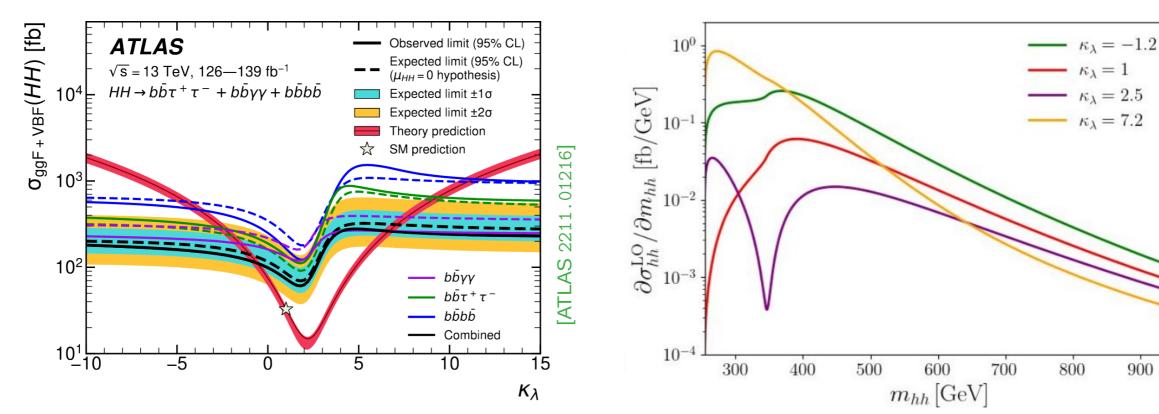


Coupling modifier:

$$\kappa_{\lambda} \equiv \frac{\lambda_{hhh}}{(\lambda_{hhh}^{(0)})^{\text{SM}}}$$

Relative change in total cross-section for varying κ_{x}

Differential m_{hh} distributions for varied κ_{x}



Note: impact of change in top Yukawa → overall shift (up/down) of distribution

900

Serdula

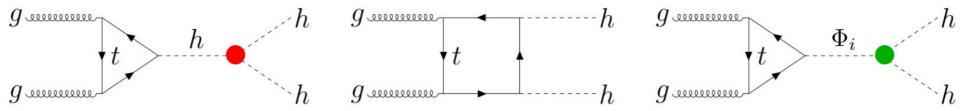
Radchenko

Plots by [K.

Di-Higgs production in arbitrary models: anyHH

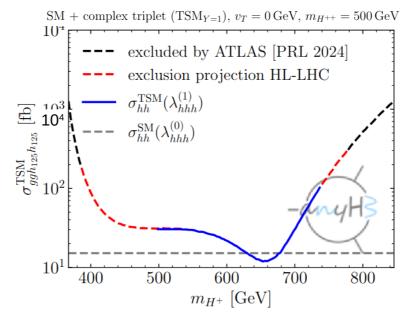
[Bahl, JB, Gabelmann, Radchenko Serdula, Weiglein *WIP*]

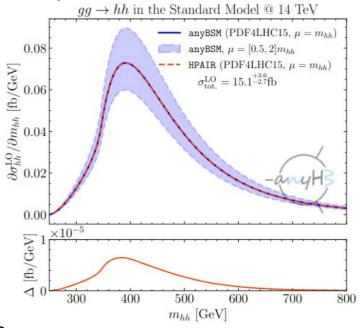
anyHH: Total and differential cross-sections (so far, at LO in QCD*) for gg → hh including 1L corrections to λ_{ijk} (computed by anyH3 [Bahl, JB, Gabelmann, Weiglein '23]) and BSM contributions and momentum-dependence in s-channel diagrams



- > Takes UFO model files as inputs, as anyH3. So far limited to models without additional coloured particles.
- Here: example results for the total di-Higgs cross-section in a model with an additional complex triplet.

Left: total cross-section vs triplet mass in triplet extension of SM Right: differential cross-section in SM, compared with HPAIR + with uncertainty band from factorisation scale in PDFs





- > Other approach: calculations in EFTs, e.g. HEFT, including higher-order QCD corrections
 - → see e.g. [Buchalla et al. '18], [Heinrich et al. '20], [Bagnaschi et al. '23]

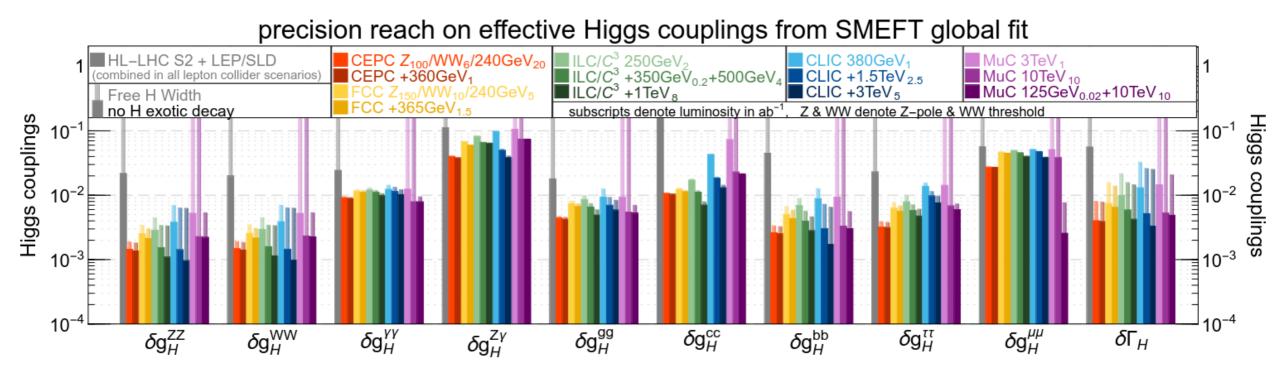
Future prospects

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Future projections for Higgs coupling measurements

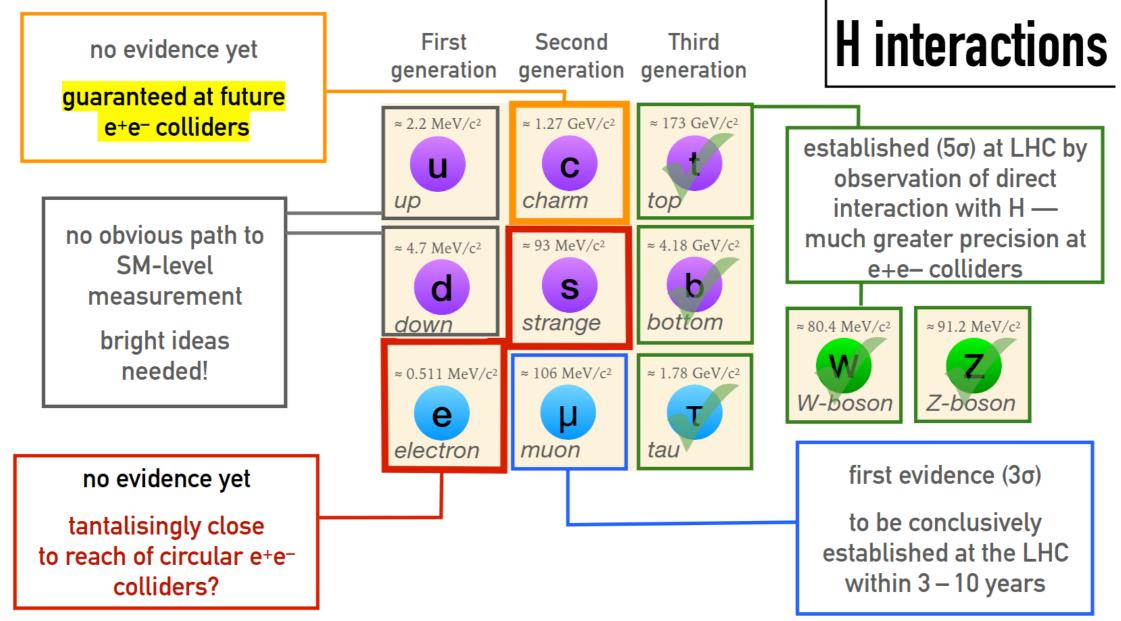
Global fit in SMEFT, using Higgs data, EW precision observables, di-boson data

e.g. [Snowmass Higgs topical report '22]



- → important to properly assess and compare prospects at future colliders
- \rightarrow keep in mind: these numbers also **depend on theoretical uncertainties** (e.g. on calculation of relevant cross-sections) \rightarrow need to be taken into account and estimated realistically!

Future prospects for Higgs coupling measurements



Direct probes of λ_{hhh} at e^+e^- colliders

- > Double-Higgs production, either in e⁺e⁻→Zhh or e⁺e⁻→vvhh
- Relies however on being above the Zhh threshold!

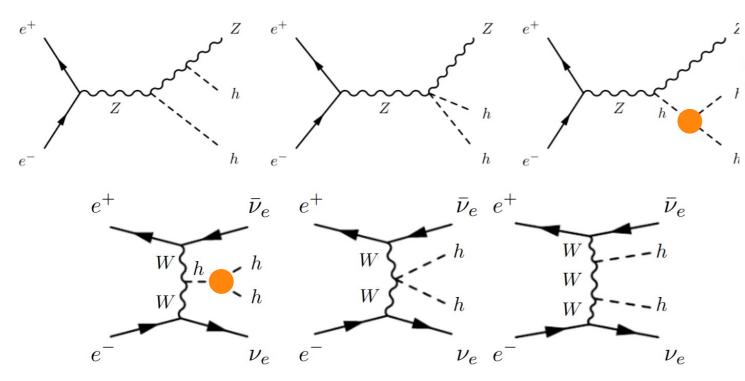
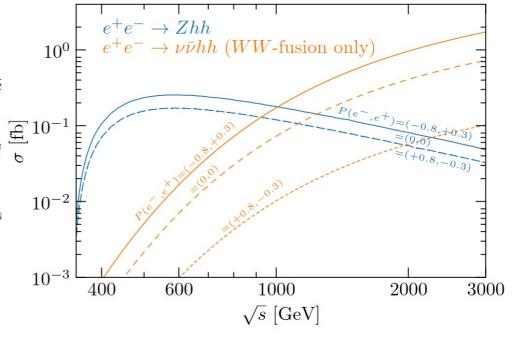


Figure from [De Blas et al. 1905.03764]

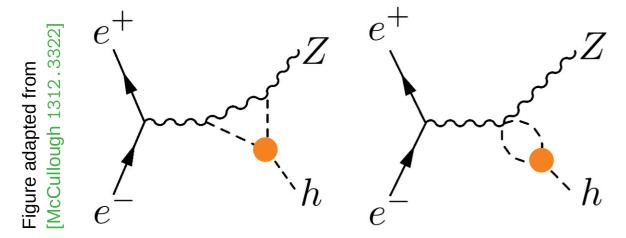
Figure from [De Blas et al. 1812.02093]



- e⁺e⁻→Zhh better at √s~500 GeV
- $^{\flat}$ e⁺e⁻ → ννhh better for larger √s

Indirect probes of λ_{hhh} at e⁺e⁻ colliders

- > Below the Zhh threshold, λ_{hhh} can still be investigated through its (indirect) effect in quantum corrections to single-Higgs production
- In particular, λ_{hhh} enters NLO corrections to e⁺e⁻ → Zh
 First pointed out in [McCullough '13], numerous works since (also with global analyses in EFT setting)



Reliable extraction of λ_{hhh} requires a consistent theory framework and control of e⁺e⁻ → Zh calculation (including e.g. effects of other BSM operators, etc.) → work in progress

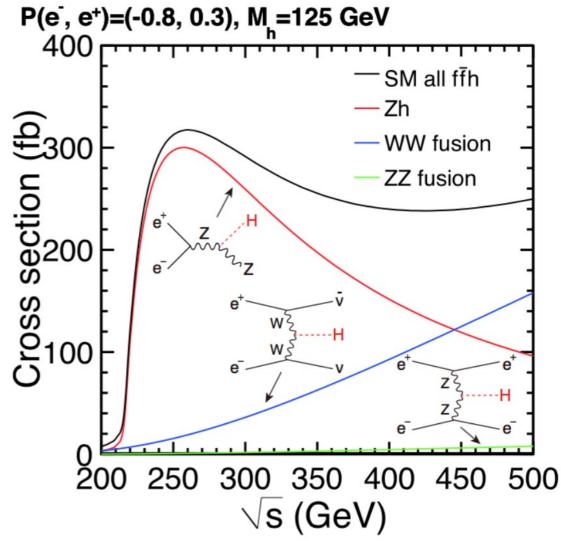
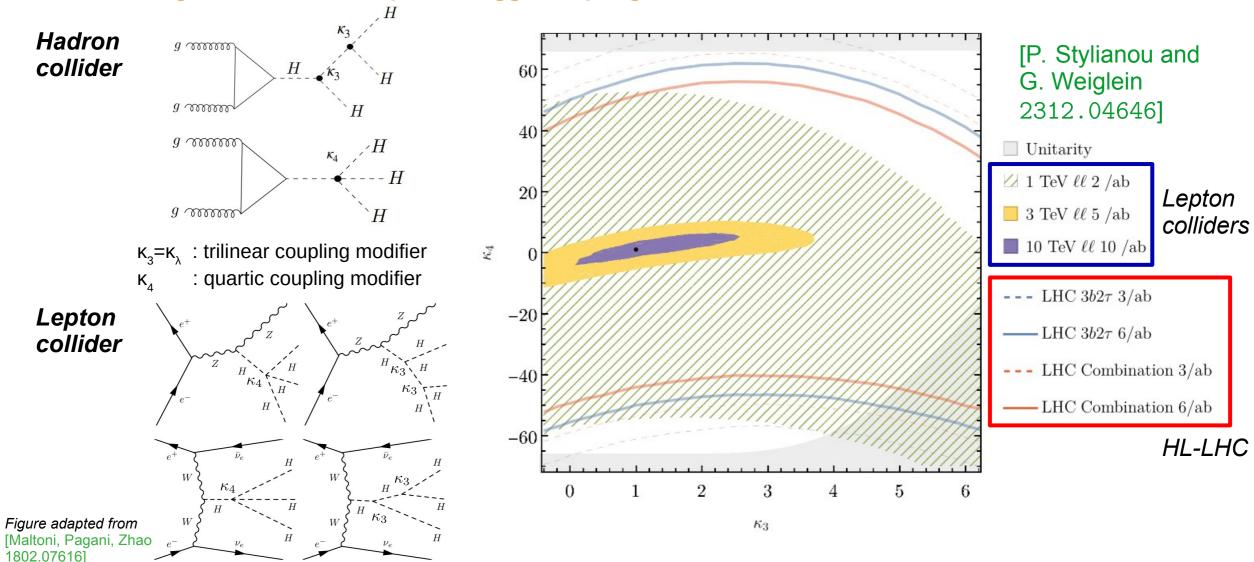


Figure from [Fujii et al. 1710.07621]

New investigations via triple-Higgs production

Constraining the trilinear and quartic Higgs couplings at the same time



Summary

- Detected Higgs boson, h₁₂₅, plays a central role in investigating the Nature of Physics Beyond the Standard Model
- Exciting times ahead, with **precision measurements of Higgs boson properties** ongoing at LHC and to be continued at future colliders (HL-LHC, e⁺e⁻ colliders, etc.)
- High-precision theory predictions are crucial to properly interpret experimental data in terms of potential discovery, or constraints on the allowed parameter space of New Physics
- Active efforts underway to improve theory calculations in SM and variety of BSM models, with also a push towards automation

Thank you very much for your attention!

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Addendum 1: The need for a Higgs boson

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Masses of elementary particles

- > Strong, weak and electromagnetic fundamental interactions described as gauge theories
 - Quantum Chromodynamics (QCD) → SU(3)_c
 - Electroweak (EW) interactions \rightarrow SU(2)_L x U(1)_Y
- ➤ Underlying gauge theories is the principle of **gauge invariance**, which strongly constrains allowed terms in the Lagrangian.

For instance, under a finite local transformation V(x) of a gauge group G, a gauge field A_{μ} transforms as

$$A_{\mu} \xrightarrow{V \in G} VAV^{-1} + \frac{i}{q}V(\partial_{\mu}V^{\dagger})$$

thus a mass term $m_A^2 A_\mu A^\mu$ is forbidden by gauge invariance

➤ Additionally, the currently-known fermions are **chiral**, i.e. weak interactions treat left-handed and right-handed fermions differently → **mass terms for chiral fermions are also forbidden by gauge invariance** e.g.

$$m_e \bar{e}_L e_R + \mathrm{h.c.}$$

 $Y=+1$ $Y=-2$
& part of $SU(2)_L$ doublet & part of $SU(2)_L$ singlet

Remember: $\bar{\psi}=\psi^{\dagger}\begin{pmatrix}0&1\\1&0\end{pmatrix}$

- > How can we explain the **observed masses** of EW gauge bosons and fermions?
 - → Brout-Englert-Higgs mechanism

Brout-Englert-Higgs mechanism

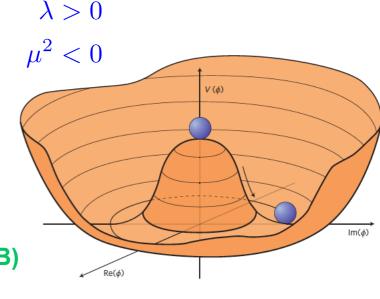
▶ Idea (in its minimal realisation): introduce a scalar* Φ – the Higgs field – doublet under SU(2)_L and with hypercharge Y=+1, and with potential

$$V(\Phi) = \mu^2 |\Phi|^2 + \lambda |\Phi|^4$$

* Why a scalar? → so that it can get a vacuum expectation value without breaking Lorentz symmetry

- \triangleright The potential V(Φ) itself (and thus also the Lagrangian of the theory) obeys the fundamental SU(2)_L x U(1)_Y gauge symmetry but the **vacuum does not**
- In other words, the Higgs field acquires a non-zero vacuum expectation value v that triggers the spontaneous breaking of the EW symmetry (EWSB)

$$\Phi = \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix} \longrightarrow \langle \Phi \rangle = \begin{pmatrix} 0 \\ v/\sqrt{2} \end{pmatrix}$$



➤ Vacuum remains symmetric under **U(1)**_{QED} gauge group (otherwise there would be charge breaking with strong phenomenological consequences!)

$$SU(2)_L \times U(1)_Y \xrightarrow{\text{EWSB}} U(1)_{\text{QED}}$$

Brout-Englert-Higgs mechanism and particle masses

$$V(\Phi) = \mu^2 |\Phi|^2 + \lambda |\Phi|^4 \quad \lambda > 0 \quad \mu^2 < 0$$

$$SU(2)_L \times U(1)_Y \xrightarrow{\text{EWSB}} U(1)_{\text{QED}}$$

Masses of gauge bosons via scalar kinetic term, with covariant derivative

$$D_{\mu}\Phi = \partial_{\mu}\Phi - \frac{1}{2}i \begin{pmatrix} g_2W_{\mu}^3 + g_YB_{\mu} & \sqrt{2}g_2W_{\mu}^+ \\ \sqrt{2}g_2W_{\mu}^- & -g_2W_{\mu}^3 + g_YB_{\mu} \end{pmatrix} \Phi$$
 with $\Phi = \begin{pmatrix} G^+ \\ \frac{1}{\sqrt{2}}(v+h+iG^0) \end{pmatrix}$ h: Higgs boson and $W_{\mu}^{\pm} = \frac{1}{\sqrt{2}}(W_{\mu}^1 \mp iW_{\mu}^2)$ which gives $|D_{\mu}\Phi|^2 \supset \frac{1}{4}g_2^2v^2W_{\mu}^+W^{-\mu} + \frac{1}{4}(g_2^2 + g_Y^2)v^2Z_{\mu}Z^{\mu}$ (\blacksquare) where $Z_{\mu} = \frac{g_2W_{\mu}^3 - g_YB_{\mu}}{\sqrt{g_2^2 + g_Y^2}}$

Before EWSB:

 $\Phi \rightarrow$ 4 degrees of freedom (d.o.f.) + 4 massless gauge bosons of SU(2)_L x U(1)_Y (W_1 , W_2 , W_3 , B) \rightarrow 4x2=8 d.o.f.

 \blacktriangleright After EWSB: would-be Goldstone bosons are "eaten" by gauge bosons which become massive h → 1 d.o.f + 3 massive gauge bosons W[±], Z → 3x3=9 d.o.f + 1 massless photon A → 2 d.o.f.

Exercise: rederive equation (\blacktriangle) + find the expression of the photon A in terms of W_3 and B

Brout-Englert-Higgs mechanism and particle masses

$$V(\Phi) = \mu^2 |\Phi|^2 + \lambda |\Phi|^4 \quad \lambda > 0 \quad \mu^2 < 0$$

$$SU(2)_L \times U(1)_Y \xrightarrow{\text{EWSB}} U(1)_{\text{QED}}$$

Masses of gauge bosons via scalar kinetic term, with covariant derivative

$$D_{\mu}\Phi = \partial_{\mu}\Phi - \frac{1}{2}i \begin{pmatrix} g_{2}W_{\mu}^{3} + g_{Y}B_{\mu} & \sqrt{2}g_{2}W_{\mu}^{+} \\ \sqrt{2}g_{2}W_{\mu}^{-} & -g_{2}W_{\mu}^{3} + g_{Y}B_{\mu} \end{pmatrix} \Phi$$
 with $\Phi = \begin{pmatrix} G^{+} \\ \frac{1}{\sqrt{2}}(v+h+iG^{0}) \end{pmatrix}$ h: Higgs boson and $W_{\mu}^{\pm} = \frac{1}{\sqrt{2}}(W_{\mu}^{1} \mp iW_{\mu}^{2})$ which gives $|D_{\mu}\Phi|^{2} \supset \frac{1}{4}g_{2}^{2}v^{2}W_{\mu}^{+}W^{-\mu} + \frac{1}{4}(g_{2}^{2} + g_{Y}^{2})v^{2}Z_{\mu}Z^{\mu}$ (\blacksquare) where $Z_{\mu} = \frac{g_{2}W_{\mu}^{3} - g_{Y}B_{\mu}}{\sqrt{g_{2}^{2} + g_{Y}^{2}}}$

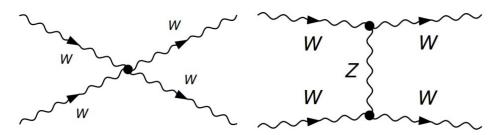
Masses of fermions (e.g. electron) via Yukawa-interaction term

$$\mathcal{L}\supset -y_e\bar{L}_L\Phi e_R + \text{h.c.} \xrightarrow{\text{EWSB}} -\frac{y_e}{\sqrt{2}}v\bar{e}_Le_R + \text{h.c.}$$

$$\xrightarrow{Y=+1}_{\substack{\text{conjugate of}\\\text{SU(2)}_{\text{L}}\text{ doublet}}} \xrightarrow{Y=-2}_{\substack{\text{SU(2)}_{\text{L}}\text{ singlet}}} v\bar{e}_Le_R + \text{h.c.}$$

Where to find "the" Higgs boson? A unitarity argument

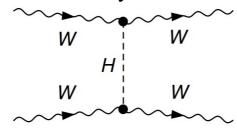
- Higgs-less alternatives to BEH mechanism were also devised (e.g. technicolor)
 - → How to **test** the BEH mechanism? At what scale can the Higgs boson be found?
- \triangleright Consider a massive boson W_u with momentum k^{μ} = (E,0,0,k)
 - \rightarrow 3 possible polarisations such that $k_{\parallel} \cdot \epsilon^{\mu} = 0$ and $\epsilon_{\parallel} \cdot \epsilon^{\mu} = -1$
 - \rightarrow 2 transverse polarisations $\varepsilon_{T_1}^{\mu} = (0,1,0,0), \ \varepsilon_{T_2}^{\mu} = (0,0,1,0)$
 - + 1 longitudinal polarisation $\varepsilon_{l}^{\mu} = (k/M_{W}, 0, 0, E/M_{W}) \sim k^{\mu}/M_{W}$ for E>>M_W
- Consider the 2→2 scattering of longitudinally polarised W bosons W, W, → W, W,
 - → without a Higgs boson, only gauge-boson diagrams like



$$\mathcal{A} \sim g_2^2 \frac{E^2}{M_W^2}$$

 ${\cal A} \sim g_2^2 {E^2 \over M_{
m W}^2}$ Loss of unitarity for large E (from ~M $_{
m W}$ /g $_{
m 2}$)!

→ adding a Higgs boson in the theory:



$$\mathcal{A}_h \sim -g_2^2 \frac{E^2}{M_W^2}$$

$$\Rightarrow \mathcal{A}_{\rm tot} \sim g_2^2 \frac{M_h^2}{M_W^2}$$
 A Higgs boson unitarises the theory if its mass < ~1 TeV

Where to find "the" Higgs boson? A unitarity argument

- Higgs-less alternatives to BEH mechanism were also devised (e.g. technicolor)
 - → How to test the BEH mechanism? At what scale can the Higgs boson be found?
- \triangleright Consider a massive W boson W₁₁ with momentum $k^{\mu} = (E,0,0,k)$



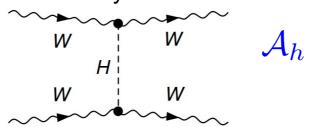
- → either a Higgs boson exists below/around the TeV scale, to unitarise gauge boson scattering in EW gauge theory or
- → some new strong dynamics would appear at ~ TeV scale

In other words, theory guaranteed that the LHC would see something!

→ adding a Higgs boson in the theory:

> Conside

 \rightarrow withd



$$\mathcal{A}_h \sim -g_2^2 \frac{E^2}{M_W^2}$$

$$\Rightarrow \mathcal{A}_{\rm tot} \sim g_2^2 \frac{M_h^2}{M_W^2}$$
 A Higgs boson unitarises the theory if its mass < ~1 TeV

Addendum 2:

Higgs and BSM

- i) Hierarchy problems
- ii) Electroweak baryogenesis
- iii) Higgs portal to dark matter
- iv) Higgs inflation
- v) Neutrino mass models with extended Higgs sectors

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Naturalness and the gauge hierarchy problem

- ➤ The EW scale is around m_{EW}^{\sim} 100 GeV (v=246 GeV) while the Planck scale, at which effects of quantum gravity must manifest themselves is M_{Pl}^{\sim} 10¹⁹ GeV \rightarrow why are there 17 orders of magnitude between m_{EW} and M_{Pl}^{\sim} ? \rightarrow (gauge) hierarchy problem
- At a more concrete level, the Higgs mass also poses a technical problem, as it is not **protected** from large (quadratic) corrections unlike for fermions and gauge bosons, nothing forbids scalar mass terms
- \blacktriangleright Let's consider the effect of a heavy BSM fermion ψ , of mass M ~ M_{\rm pl} with a Lagrangian

$${\cal L} \supset ar{\psi} (i \gamma^{\mu} \partial_{\mu} - M) \psi - y_{\psi} ar{\psi} \psi h$$

and let's compute the leading corrections to the Higgs mass in this toy model

$$\Delta^{(1\ell)} m_h^2 = -(-iy_\psi)^2 \int \frac{d^d k}{i(2\pi^2)} \operatorname{tr} \left[\frac{i(\not k + M)}{k^2 - M^2} \frac{i(\not p - \not k) + M)}{(p - k)^2 - M^2} \right]$$

$$\approx -\frac{y_\psi^2}{4\pi^2} M_{\rm Pl}^2 \quad \text{with } p^2 \ll M^2 \quad \& \ Q = M \approx M_{\rm Pl}$$

➤ Getting the Higgs mass right at 125 GeV would imply a tuning between tree-level mass and loop corrections to 32 digits!!! → technical hierarchy problem

Solutions to the gauge hierarchy problem: Supersymmetry

- Supersymmetry (SUSY): [Wess, Zumino '74] and many more Extend space-time symmetry (Poincaré group) by introducing new symmetry between fermions and bosons (SUSY is only option to circumvent Coleman-Mandula theorem [Coleman, Mandula '67], see [Haag, Lopuszanski, Sohnius '75])
 - ightarrow Each fermion (boson) has a bosonic (fermionic) superpartner, with same mass and related couplings, e.g. for toy model of previous slide, ψ has a superpartner $\tilde{\psi}$, with interaction terms

$$\mathcal{L}\supset -y_{\psi}\bar{\psi}\psi h - \lambda_{\tilde{\psi}}\tilde{\psi}^*\tilde{\psi}h^2 \quad \text{ with } \lambda_{\tilde{\psi}}=y_{\psi}^2, \ m_{\tilde{\psi}}=M_{\Psi}=M_{\Pi}$$
 such that
$$+\frac{y_{\psi}^2}{4\pi^2}M_{\Pi}^2 + \cdots + \frac{y_{\psi}^2}{4\pi^2}M_{\Pi}^2 + \cdots + \cdots + \cdots + \cdots = 0$$

- SUSY must be broken, otherwise selectron would have mass 511 keV and would have had to be seen already
- > But SUSY can be broken (super)softly, i.e. without reintroducing quadratic divergences in m,

$$m_{\tilde{\psi}}^2 = M_{\psi}^2 + \Delta m^2$$
 and $\lambda_{\tilde{\psi}} = y_{\psi}^2$

Numerous phenomenological models, such as Minimal Supersymmetric Standard Model (MSSM), Next-to-MSSM (NMSSM), Dirac gaugino models, etc., however so far no sign of SUSY at the LHC...

Solutions to the gauge hierarchy problem: Compositeness

- Compositeness: see e.g. [Agashe, Contino, Pomarol '04], [Giudice, Grojean, Pomarol, Rattazzi '07] + refs therein Light scalars already known in Nature, e.g. pions, but these are *not fundamental*, rather bound or in other words composite states
 - ightarrow Introduce a new strongly coupled sector, with a global symmetry group G, spontaneous broken down to H at a scale f $G \xrightarrow{SSB} H \supset SU(2)_L \times U(1)_Y$ NB: only a part of H is gauged!

→ Higgs boson appears as a pseudo-Goldstone boson → naturally light

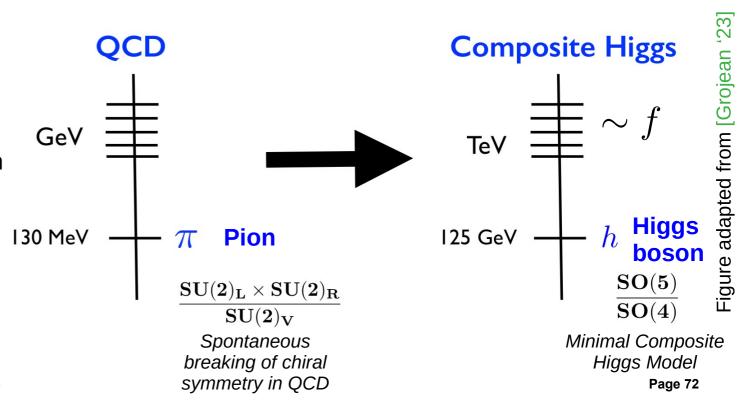
Minimal model (1 Higgs doublet):

$$\rightarrow$$
 G = SO(5) (10d); H = SO(4) (6d)

Composite Two-Higgs-Doublet Model:

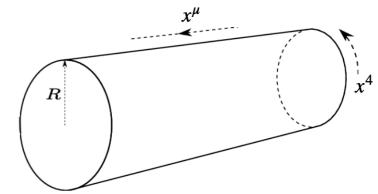
$$\rightarrow$$
 G = SO(6) (15d); H = SO(4) x SO(2) (7d)

- ➤ Ratio v/f determined by *misalignment* between directions of G/H and SU(2)_LxU(1)_Y/U(1)_{QED} breakings
- Partial compositeness to explain quark mass paterns



Other solutions to the gauge hierarchy problem

- Large Extra-dimensions: [Arkani-Hamed, Dimopoulos, Dvali '98] (see e.g. Randall-Sundrum models, [Randall, Sundrum '99])
 Add at least one more dimension of space-time, which is compactified
 - → tower of excited states (Kaluza-Klein modes)
 - + effective Planck scale in 4d is lowered

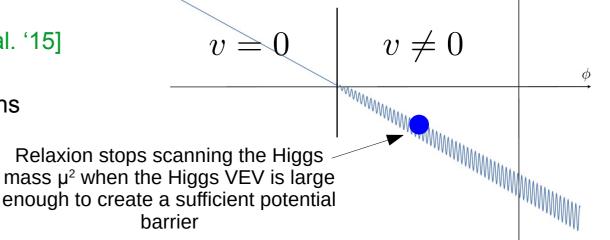


- ➤ Gauge-Higgs unification: [Manton '79], [Fairlie '79], [Hosotani '83], etc.
 - Hosotani mechanism: In 5d, a gauge boson contains 5 components
 - → 4 components = 4d gauge boson + 1 component = 4d Higgs boson (which triggers EWSB)
 - → Higgs mass is then again protected by gauge symmetry in 5d
- Cosmological relaxation:

see e.g. [Graham, Kaplan, Rajendran '15], [Espinosa et al. '15]

Promote the Higgs mass term μ^2 to a dynamical field, the **relaxion**, and give this field a potential and interactions with the Higgs boson (and VEV) such that it selects the appropriate value of μ^2

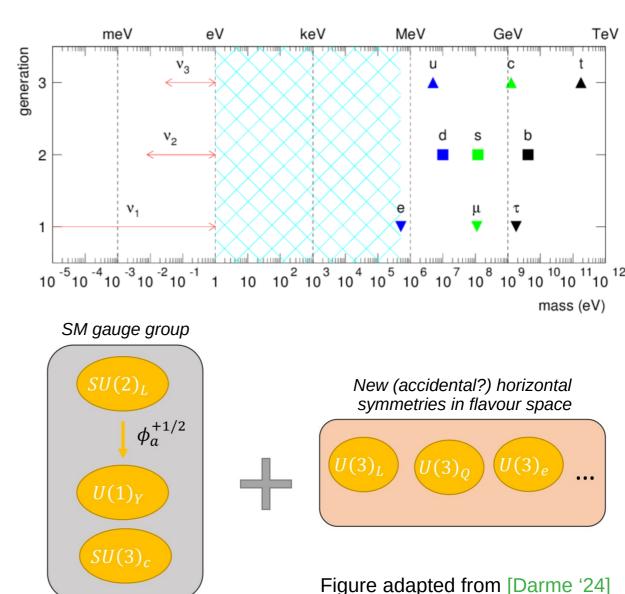
and many more...



 $V(\phi)$

The Yukawa hierarchy problem and flavour

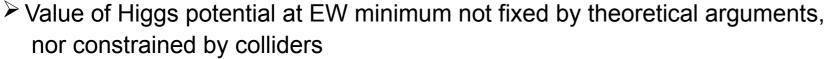
- Fermion mass patterns completely unexplained why is m_t ~ 3 x 10⁵ m_e ? (not to mention neutrinos...)
- Fermion masses in SM → entirely determined by Yukawa couplings between fermions and Higgs boson
 - → why does the Higgs treat the three fermion families (identical w.r.t gauge symmetries) so differently?
- ➤ No guiding principle in Yukawa interactions in SM
- Gauge symmetries act on all three fermion families in the same way → something must treat the families differently → for instance a "horizontal symmetry" ?



The cosmological constant and its fine-tuning problem

- ➤ Cosmological observations → Universe expanding at accelerating pace
- Explained in ΛCDM model by cosmological constant, corresponding to a vacuum energy:

[Planck '15]
$$\rho_{vac} \sim 2.5 \times 10^{-47} \text{ GeV}^4$$



$$V(\Phi) = \mu^{2} |\Phi|^{2} + \lambda |\Phi|^{4} + V_{0}$$

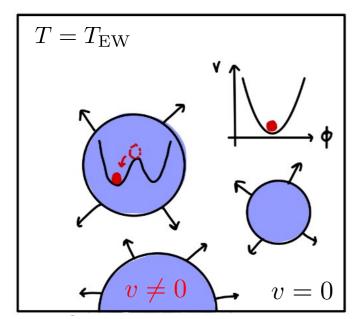
$$\longrightarrow V_{\min} = \frac{1}{2} \mu^{2} v^{2} + \frac{1}{4} \lambda v^{4} + V_{0} = -1.2 \times 10^{8} \text{ GeV}^{4} + V_{0} = \rho_{\text{vac}} \sim 2.5 \times 10^{-47} \text{ GeV}^{4}$$

- \triangleright Cancellation/fine-tuning of \sim 55 digits needed in V_0 to reproduce the measured vacuum energy!
 - → cosmological constant problem
- Possible solutions involve anthropic principle (multiverse), modifications of GR/ΛCDM, or of QFT, etc.

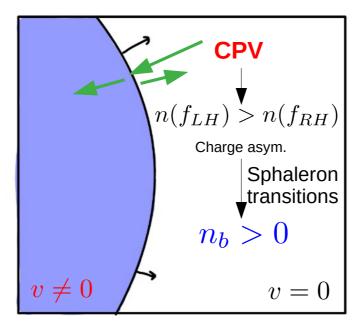
Electroweak Baryogenesis – a brief sketch

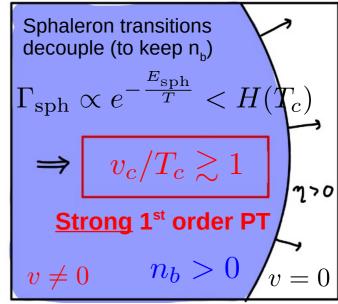
- Sakharov conditions in EWBG
 - 1) Baryon number violation
 - 2) C and CP violation
 - 3) Loss of thermal equilibrium

- → Sphaleron transitions (break B+L)
- → C violation + CP violation in extended Higgs sector
- → Loss of th. eq. via a strong 1st order EWPT



1) Bubble nucleation





2) Baryon number generation 3) Baryon number conservation

EWBG only involves phenomena around the EW scale \rightarrow testable in the foreseeable future via λ_{hhh} , collider searches, gravitational waves or primordial black holes (sourced by 1st order EWPT)

=igure adapted from [Biermann '22]

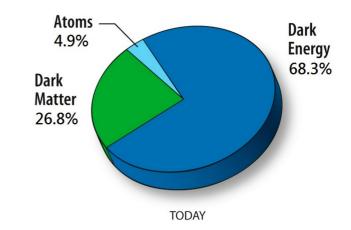
Higgs portal to dark sectors

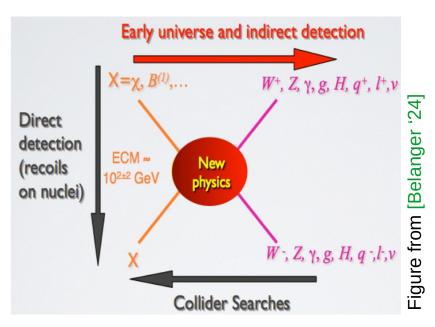
➤ Dark matter (DM)

- Non-relativistic matter (→ can't be neutrinos)
- Only/mostly gravitational interactions → several types of astrophysical evidence (e.g. galaxy rotation curves, etc.)
- · Collisionless (c.f. Bullet cluster) & pressureless
- Needed to seed large-structure formation
- → No SM particle can fit this!
- |Φ|² is a gauge singlet → Higgs field provides a perfect way to write a portal term in the Lagrangian,
 e.g. simplest example = add to SM a singlet S, charged under a global Z₂ symmetry to stabilise DM

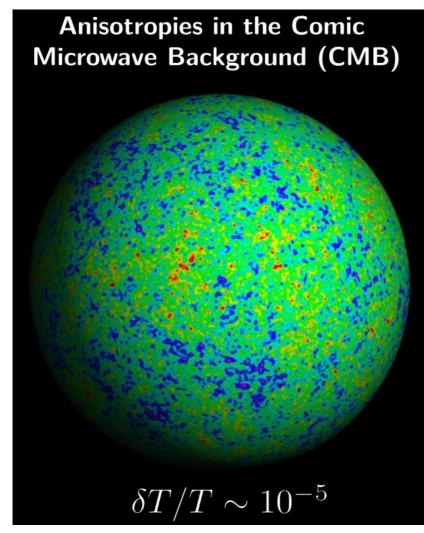
$$\mathcal{L}_{\mathbb{Z}_2 ext{SSM}} = \mathcal{L}_{ ext{SM}} - \lambda_{ ext{portal}} S^2 |\Phi|^2 - \lambda_{ ext{dark}} S^4$$
 $\lambda_{ ext{portal}}$: controls DM relic density & detection

➤ Plethora of models: inert singlets, doublets, triplets; Next-to-Two-Higgs-Doublet Model (N2HDM), S2HDM, etc.

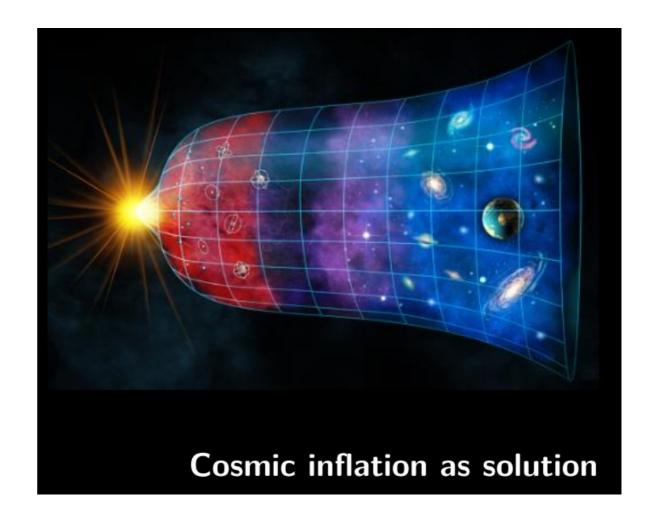




Cosmic inflation



[Planck '18]



- Phase of exponential growth driven by scalar field inflaton with very flat potential → slow-roll inflation
- What if the Higgs boson plays the role of the inflaton?
 [Bezrukov, Shaposhnikov '07]
 - → Higgs inflation
 - \rightarrow Higgs coupled **non-minimally** to gravity

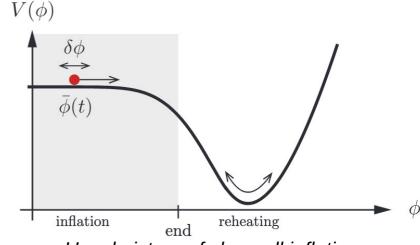
$$\mathcal{L} = \mathcal{L}_{\mathrm{SM}} - rac{1}{2} M_{\mathrm{Pl}}^2 R - \xi |\Phi|^2 R$$
 (in Jordan frame)

Change from *Jordan frame* (in which Lagrangian is written) to *Einstein frame* (with canonical coupling to gravity)

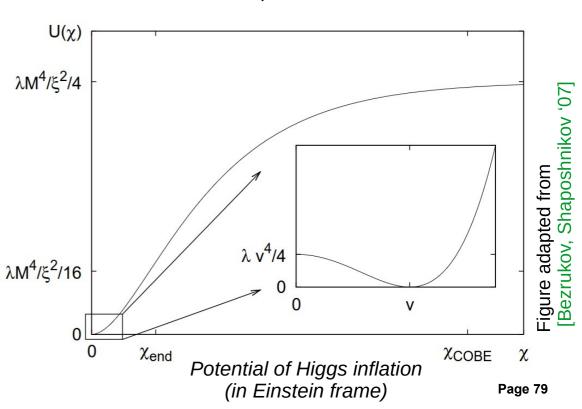
$$g^E_{\mu\nu} = \Omega^2(h)g^J_{\mu\nu}, \qquad \text{with } \Omega^2(h) = 1 + \frac{\xi h^2}{M_{\rm Pl}^2}$$

$$\Rightarrow \mathcal{L}^E \supset -\frac{1}{2}M_{\rm Pl}^2R^E + \frac{1}{2}(\partial_\mu\chi)^2 - \underbrace{\frac{\lambda}{4\Omega^4(h(\chi))}\big(h(\chi)^2 - v^2\big)^2}_{\Xi U(\chi)}$$
 X: Higgs field in Einstein frame

➤ Numerous developments (non-minimal Higgs sectors, different couplings, etc.)



Usual picture of slow-roll inflation



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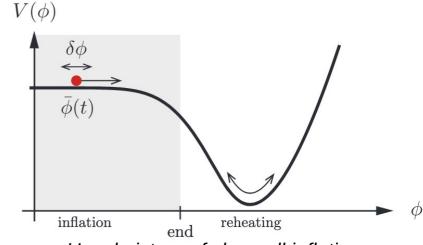
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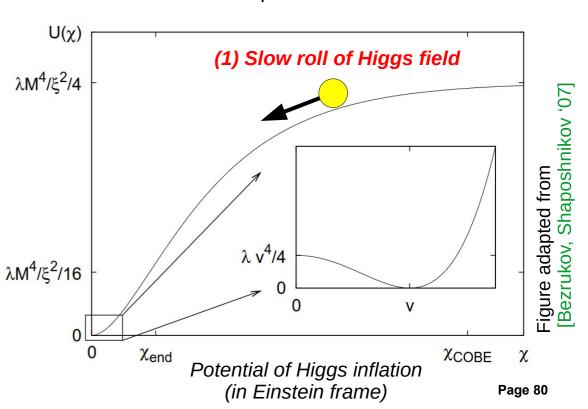
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$$\equiv U(\chi)$$

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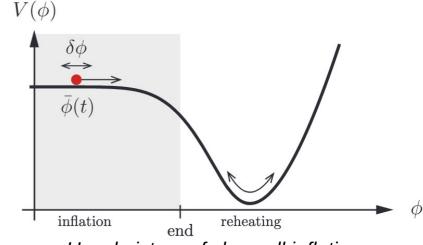
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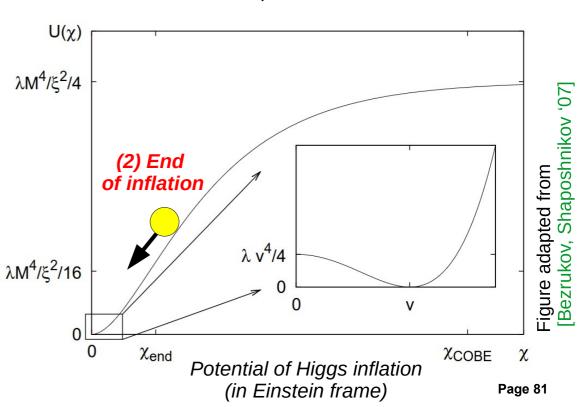
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$$\Rightarrow \mathcal{L}^E \supset -\frac{1}{2}M_{\rm Pl}^2R^E + \frac{1}{2}(\partial_\mu\chi)^2 - \underbrace{\frac{\lambda}{4\Omega^4(h(\chi))}\big(h(\chi)^2 - v^2\big)^2}_{\Xi U(\chi)}$$
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Usual picture of slow-roll inflation



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 (in Jordan frame)

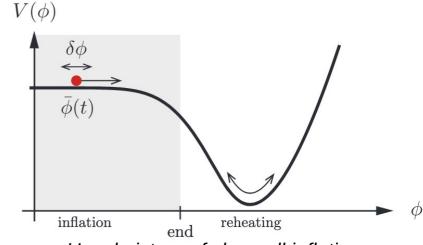
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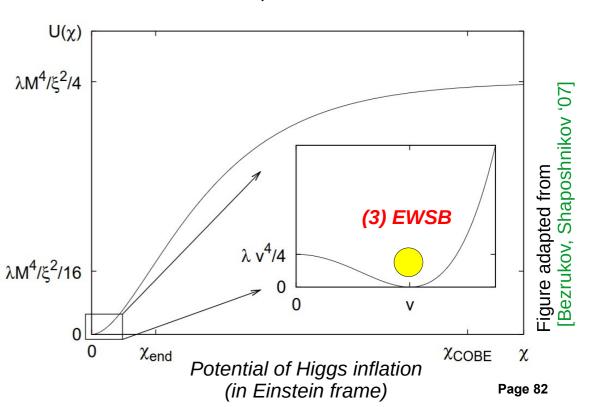
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$$\equiv U(\chi)$$

Numerous developments (non-minimal Higgs sectors, different couplings, etc.)



Usual picture of slow-roll inflation



- However, since 1960's early signs of neutrino oscillations ("solar neutrino deficit"), eventually confirmed ~25 years ago
 - → atmospheric neutrino oscillations in 1998
 - → solar neutrino oscillations in 2001
 - → 2015 Nobel Prize for Kajita and McDonald



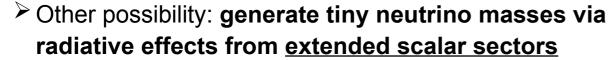
→ neutrinos do have masses → extension of SM needed!



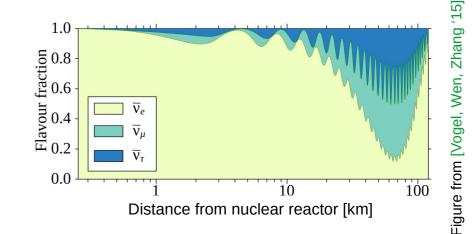
→ basic idea (type I): introduce, heavy, right-handed Majorana neutrinos (RHN) N_R

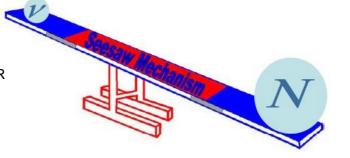
$$\Rightarrow m_{\nu_L} \sim \frac{y_\nu^2 v^2}{M_R}$$

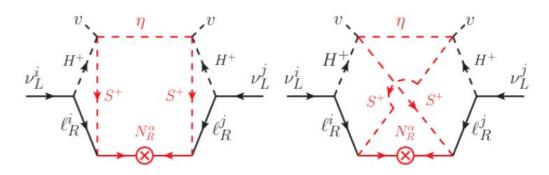
experimentally



- \rightarrow [Zee '80], [Babu '88], [Aoki, Kanemura, Seto '08], etc.
- → no longer need for very heavy RHN







An example of radiative neutrino mass generation: the **Aoki-Kanemura-Seto model** Figure from [Aoki, Enomoto, Kanemura '22] Page 83

Addendum 3: $M_{\rm w}$ calculations in the SM and beyond

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M_w calculation in the SM I

See e.g. [Awramik, Czakon, Freitas, Weiglein '03], [Hessenberger TUM thesis '18]

- Base for MW calculation is the decay of the muon
 - \succ Extract G_F from muon lifetime T_u by computing T_u in the Fermi theory

$$\frac{1}{\tau_{\mu}} = \frac{G_F^2 m_{\mu}^5}{192\pi^3} F(m_e^2/m_{\mu}^2) \bigg(1 + \frac{3}{5} \frac{m_{\mu}^2}{M_W^2}\bigg) (1 + \Delta q)$$
 with $F(x) \equiv 1 - 8x - 12x^2 \ln x + 8x^3 - x^4$
$$Tree-level \ W \ propagator \ contributions \ (not \ in \ Fermi \ th. \ but \ numerically \ tiny)$$

> Relate M_w, M_z, α, G_F by computing muon decay in SM, and matching to Fermi theory result

$$\frac{G_F}{\sqrt{2}} = \frac{e^2}{8M_W^2 s_W^2} (1+\Delta r) \quad \Rightarrow \quad M_W^2 \left(1-\frac{M_W^2}{M_Z^2}\right) = \frac{\pi\alpha}{\sqrt{2}G_F} (1+\Delta r) \qquad \text{OS scheme}$$

 $\Delta r \equiv \Delta r(M_w, M_z, m_h, m_t, ...)$ denotes corrections to muon decay (w/o finite QED effects)

 \rightarrow Previous relation used to determine M_w as solution, via iterations, of

$$M_W^2 = M_Z^2 \left[\frac{1}{2} + \sqrt{\frac{1}{4} - \frac{\pi \alpha}{\sqrt{2} G_F M_Z^2}} \left(1 + \Delta r(M_W^2, M_Z^2, m_h^2, m_t^2, \cdots) \right) \right] \qquad \text{OS scheme}$$

M_w calculation in the SM II

$$\frac{G_F}{\sqrt{2}} = \frac{e^2}{8M_W^2 s_W^2} (1 + \Delta r) \quad \Rightarrow \quad M_W^2 \left(1 - \frac{M_W^2}{M_Z^2} \right) = \frac{\pi \alpha}{\sqrt{2} G_F} (1 + \Delta r)$$

$$\frac{G_F}{\sqrt{2}} = \frac{e^2}{8M_W^2 s_W^2} (1 + \Delta r) \quad \Rightarrow \quad M_W^2 \left(1 - \frac{M_W^2}{M_Z^2} \right) = \frac{\pi \alpha}{\sqrt{2} G_F} (1 + \Delta r)$$

$$M_W^2 = M_Z^2 \left[\frac{1}{2} + \sqrt{\frac{1}{4} - \frac{\pi \alpha}{\sqrt{2} G_F M_Z^2}} \left(1 + \Delta r (M_W^2, M_Z^2, m_h^2, m_t^2, \cdots) \right) \right]$$

At one loop

$$\Delta r^{(1)} = 2\delta^{(1)} Z_e + \frac{\Sigma_{WW}^{(1)}(p^2 = 0) - \delta^{(1)} M_W^2}{M_W^2} - \frac{\delta^{(1)} s_W^2}{s_W^2} + \{\text{vertex + box corrections}\}$$

 Σ_{ww} : transverse part of the W-boson self-energy, $\delta^{(1)}X$: 1L counterterm to quantity X

One can show that

$$\delta^{(1)} Z_e \simeq \frac{1}{2} \Delta \alpha + \cdots \quad \text{and} \quad \frac{\delta^{(1)} s_W^2}{s_W^2} \simeq \frac{c_W^2}{s_W^2} \Delta \rho^{(1)}$$
with
$$\Delta \alpha = \frac{\partial}{\partial p^2} \Sigma_{\gamma\gamma} \big|_{p^2 = 0} - \frac{\text{Re} \Sigma_{\gamma\gamma} (p^2 = M_Z^2)}{M_Z^2}$$

Leading terms can be rewritten as [Sirlin '80]

$$\Delta r^{\alpha} = \Delta \alpha - \frac{c_W^2}{s_W^2} \Delta \rho^{(1)} + \Delta r_{\text{remainder}}(m_h)$$

with $\Delta\alpha$: contribution from light fermion loops to photon vacuum polarisation $\Delta \rho$: corrections to the ρ parameter

$$\rho \equiv \frac{G_{\text{NC}}}{G_{\text{CC}}} \quad \Rightarrow \quad \rho^{(0)} = \frac{M_W^2}{c_W^2 M_Z^2} = 1 \text{ and } \Delta \rho^{(1)} = \frac{\Sigma_{ZZ}^{(1)}(p^2 = 0)}{M_Z^2} - \frac{\Sigma_{WW}^{(1)}(p^2 = 0)}{M_W^2}$$

M_w calculation in the SM III

$$\frac{G_F}{\sqrt{2}} = \frac{e^2}{8M_W^2 s_W^2} (1 + \Delta r) \quad \Rightarrow \quad M_W^2 \left(1 - \frac{M_W^2}{M_Z^2} \right) = \frac{\pi \alpha}{\sqrt{2} G_F} (1 + \Delta r)$$

$$M_W^2 = M_Z^2 \left[\frac{1}{2} + \sqrt{\frac{1}{4} - \frac{\pi \alpha}{\sqrt{2} G_F M_Z^2} \left(1 + \Delta r(M_W^2, M_Z^2, m_h^2, m_t^2, \cdots) \right)} \right]$$

At higher orders

$$\Delta r = \Delta r^{\alpha} + \Delta r^{\alpha \alpha_s} + \Delta r^{\alpha \alpha_s^2} + \Delta r^{\alpha \alpha_s^3 m_t} + \Delta r_{\rm ferm}^{\alpha^2} + \Delta r_{\rm bos}^{\alpha^2} + \Delta r_{\rm ferm}^{G_F^2 \alpha_s m_t^4} + \Delta r_{\rm ferm}^{G_S^3 m_t^6} + \Delta r_{\rm bos}^{G_S^2 \alpha_s m_t^4} + \Delta r_{\rm ferm}^{G_S^3 m_t^6} + \Delta r_{\rm bos}^{G_S^3 m_t^6} + \Delta r_{\rm ferm}^{G_S^3 m_t^6} + \Delta r_{\rm bos}^{G_S^3 m_t^6} + \Delta r_{\rm bos}^{G_S m_t^6} + \Delta r_{\rm bo$$

[Awramik, Czakon, Freitas, Weiglein '03] gives a parametrisation as

$$M_{W} = M_{W}^{0} - c_{1} dH - c_{2} dH^{2} + c_{3} dH^{4} + c_{4} (dh - 1) - c_{5} d\alpha + c_{6} dt - c_{7} dt^{2} - c_{8} dH dt + c_{9} dh dt - c_{10} d\alpha_{s} + c_{11} dZ,$$

with

$$dH = \ln\left(\frac{M_{\rm H}}{100 \text{ GeV}}\right), \quad dh = \left(\frac{M_{\rm H}}{100 \text{ GeV}}\right)^2, \quad dt = \left(\frac{m_{\rm t}}{174.3 \text{ GeV}}\right)^2 - dZ = \frac{M_{\rm Z}}{91.1875 \text{ GeV}} - 1, \quad d\alpha = \frac{\Delta \alpha}{0.05907} - 1, \quad d\alpha_{\rm s} = \frac{\alpha_{\rm s}(M_{\rm Z})}{0.119} - 1,$$

$$\begin{array}{ll} \mathbf{n} \\ \mathrm{dH} = \ln \left(\frac{M_{\mathrm{H}}}{100 \; \mathrm{GeV}} \right), & \mathrm{dh} = \left(\frac{M_{\mathrm{H}}}{100 \; \mathrm{GeV}} \right)^{2}, & \mathrm{dt} = \left(\frac{m_{\mathrm{t}}}{174.3 \; \mathrm{GeV}} \right)^{2} - 1, \\ \mathrm{dZ} = \frac{M_{\mathrm{Z}}}{91.1875 \; \mathrm{GeV}} - 1, & \mathrm{d}\alpha = \frac{\Delta\alpha}{0.05907} - 1, & \mathrm{d}\alpha_{\mathrm{s}} = \frac{\alpha_{\mathrm{s}}(M_{\mathrm{Z}})}{0.119} - 1, \\ \end{array} \quad \begin{array}{ll} M_{\mathrm{W}}^{0} = 80.3779 \; \mathrm{GeV}, & c_{1} = 0.05263 \; \mathrm{GeV}, & c_{2} = 0.010239 \; \mathrm{GeV}, \\ c_{3} = 0.000954 \; \mathrm{GeV}, & c_{4} = -0.000054 \; \mathrm{GeV}, & c_{5} = 1.077 \; \mathrm{GeV}, \\ c_{6} = 0.5252 \; \mathrm{GeV}, & c_{7} = 0.0700 \; \mathrm{GeV}, & c_{8} = 0.004102 \; \mathrm{GeV}, \\ c_{9} = 0.000111 \; \mathrm{GeV}, & c_{10} = 0.0774 \; \mathrm{GeV}, & c_{11} = 115.0 \; \mathrm{GeV}, \end{array}$$

Note: Δr also serves to extract the Higgs VEV from G₋

$$v^2 = \frac{1}{\sqrt{2}G_E}(1 + \Delta r)$$

M_w calculation beyond the SM

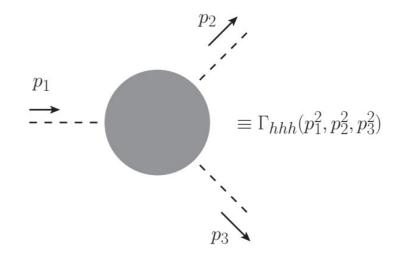
- Idea of the calculation remains the same, but full theory calculation (that is matched with the Fermi theory one) is now done in the BSM model
- > In BSM models, M_W (→ muon decay) can receive contributions both at **tree level** and at **loop level**. Considering a model with both sources (and turning to $\overline{\rm MS}$ for simplicity just here), one can write at 1L [Athron et al. 1710.03760, 2204.05285] $M_W^2 \Big|^{\overline{\rm MS}} = (M_W^{\rm SM}|^{\overline{\rm MS}})^2 \Big\{ 1 + \frac{s_W^2}{c_W^2 s_W^2} \Big[\frac{c_W^2}{s_W^2} (\Delta \rho_{\rm tree} + \Delta \rho_{\rm loop}^{\rm BSM}) \Delta r_{\rm remainder}^{\rm BSM} \Delta \alpha^{\rm BSM} \Big] \Big\}$
- In the following, we will only discuss models with $\rho^{(0)}=1$, and we stay in **OS scheme**
- Some 2L corrections to Δρ known in BSM models
 - \rightarrow O($\alpha\alpha_s$) SUSY corrections in [Djouadi et al. '96, '98]
 - $> O(\alpha_1^2, \alpha_1 \alpha_2, \alpha_2^2)$ in MSSM in [Heinemeyer, Weiglein '02], [Hastier, Heinemeyer, Stöckinger, Weiglein '05]
 - BSM scalar + top quark corrections in (aligned) 2HDM and IDM [Hessenberger, Hollik '16]
- ightharpoonup Inclusion of known higher-order SM corrections crucial $\Delta r = \Delta r^{
 m SM} + \Delta r^{
 m BSM}$
- \rightarrow Calculations of M_w with Δ r to full BSM 1L + partial BSM 2L (from resummation and Δ p) + SM up to 4L
 - MSSM [Heinemeyer, Hollik, Weiglein, Zeune '13]
 - NMSSM [Stål, Weiglein, Zeune '15]
 - MRSSM [Diessner, Weiglein '19]
 - 2HDM & IDM [Hessenberger '18] (TUM thesis and code THDM_EWPOS)

Addendum 4: Calculations of λ_{hhh}

DESY.

An effective Higgs trilinear coupling

In principle: consider 3-point function Γ_{hhh} but this is momentum dependent \rightarrow very difficult beyond one loop



Instead, consider an effective trilinear coupling

$$\lambda_{hhh} \equiv \left. \frac{\partial^3 V_{\text{eff}}}{\partial h^3} \right|_{\text{min}}$$

entering the coupling modifier

$$\kappa_{\lambda} = \frac{\lambda_{hhh}}{(\lambda_{hhh}^{(0)})^{\text{SM}}} \qquad \text{with } (\lambda_{hhh}^{(0)})^{\text{SM}} = \frac{3m_h^2}{v}$$

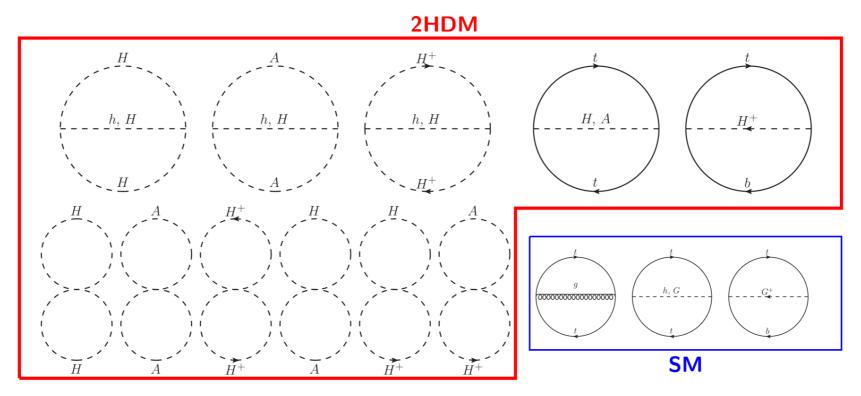
constrained by experiments (applicability of this assumption discussed later)

Effective-potential calculation

[JB, Kanemura '19]

> Step 1: compute
$$V_{\text{eff}} = V^{(0)} + \frac{1}{16\pi^2}V^{(1)} + \frac{1}{(16\pi^2)^2}V^{(2)}$$
 (MS result)

- → V⁽²⁾: 1PI vacuum bubbles
- → Dominant BSM contributions to $V^{(2)}$ = diagrams involving heavy BSM scalars and top quark
- → Neglect masses of light states (SM-like Higgs, light fermions, ...)



Effective-potential calculation

[JB, Kanemura '19]

> Step 1: compute
$$V_{\rm eff}=V^{(0)}+\frac{1}{16\pi^2}V^{(1)}+\frac{1}{(16\pi^2)^2}V^{(2)}$$
 (MS result)

- → V⁽²⁾: 1PI vacuum bubbles
- → Dominant BSM contributions to $V^{(2)}$ = diagrams involving heavy BSM scalars and top quark

> Step 2: derive an effective trilinear coupling

$$\frac{\lambda_{hhh}}{\text{(MS result too)}} \equiv \frac{\partial^3 V_{\text{eff}}}{\partial h^3} \bigg|_{\text{min.}} = \frac{3[M_h^2]_{V_{\text{eff}}}}{v} + \left[\frac{\partial^3}{\partial h^3} - \frac{3}{v} \left(\frac{\partial^2}{\partial h^2} - \frac{1}{v} \frac{\partial}{\partial h}\right)\right] \Delta V \bigg|_{\text{min.}}$$

Express tree-level result in terms of effective-potential Higgs mass

Effective-potential calculation

[JB, Kanemura '19]

> Step 1: compute
$$V_{\rm eff} = V^{(0)} + \frac{1}{16\pi^2}V^{(1)} + \frac{1}{(16\pi^2)^2}V^{(2)}$$
 (MS result)

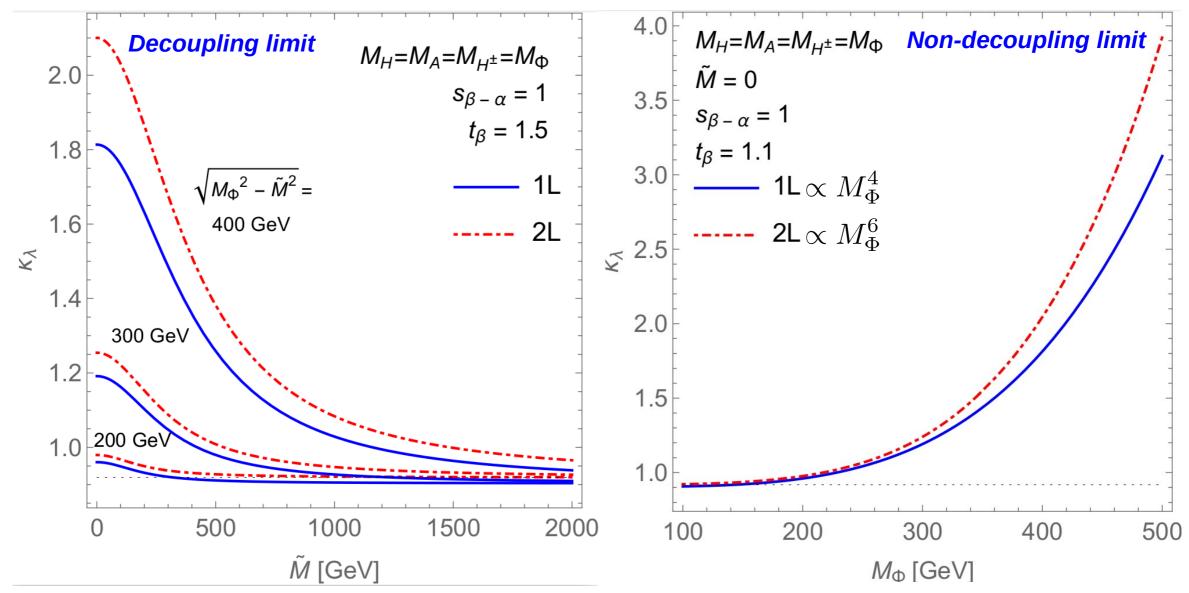
- → V⁽²⁾: 1PI vacuum bubbles
- → Dominant BSM contributions to $V^{(2)}$ = diagrams involving heavy BSM scalars and top quark

Step 2:
$$\lambda_{hhh} \equiv \left. \frac{\partial^3 V_{\rm eff}}{\partial h^3} \right|_{\rm min.} = \left. \frac{3[M_h^2]_{V_{\rm eff}}}{v} + \left[\frac{\partial^3}{\partial h^3} - \frac{3}{v} \left(\frac{\partial^2}{\partial h^2} - \frac{1}{v} \frac{\partial}{\partial h} \right) \right] \Delta V \right|_{\rm min}$$
 (MS result too)

- > **Step 3**: conversion from \overline{MS} to OS scheme
 - \Rightarrow Express result in terms of **pole masses**: M_t, M_h, M_{Φ} (Φ =H,A,H $^{\pm}$); OS Higgs VEV $v_{\rm phys} = \frac{1}{\sqrt{\sqrt{2}G_F}}$
 - o Include finite WFR: $\hat{\lambda}_{hhh} = (Z_h^{\mathrm{OS}}/Z_h^{\overline{\mathrm{MS}}})^{3/2}\lambda_{hhh}$
 - ightharpoonup Prescription for M to ensure **proper decoupling** with $M_\Phi^2 = \tilde{M}^2 + \tilde{\lambda}_\Phi v^2$ and $\tilde{M} \to \infty$

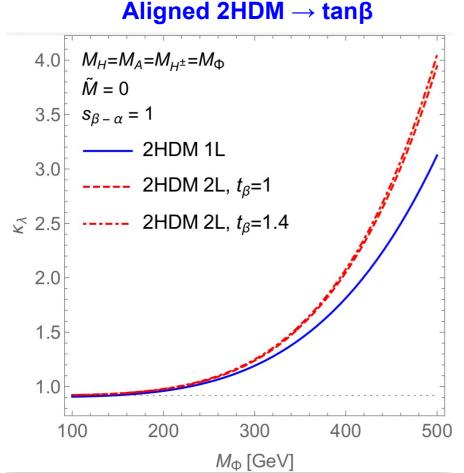
Our results in the aligned 2HDM

Taking degenerate BSM scalar masses: $M_{\phi} = M_{H} = M_{A} = M_{H}^{\pm}$

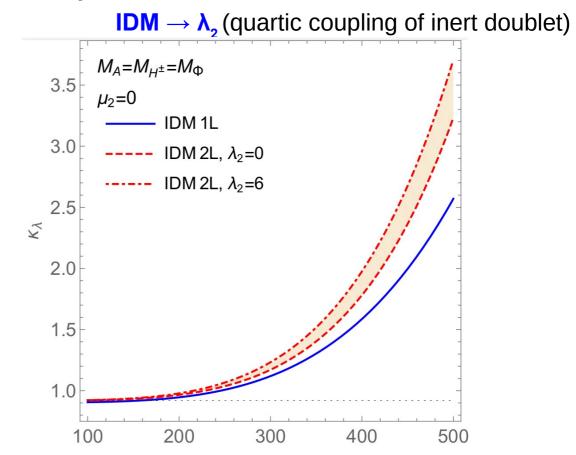


$\lambda_{\mbox{\tiny hhh}}$ at two loops in more models

- Calculations in several other models: Inert Doublet Model (IDM), singlet extension of SM
- Each model contains a **new parameter appearing from two loops**:



 $tan\beta$ constrained by perturbative unitarity \rightarrow only small effects



 λ_2 is less contrained \rightarrow enhancement is possible (but 2L effects remain <u>well smaller</u> than 1L ones)

 M_{Φ} [GeV]

Constraining BSM models with λ_{hhh}

i. Can we apply the limits on κ_{λ} , extracted from experimental searches for di-Higgs production, for BSM models?

ii. Can large BSM deviations occur for points still allowed in light of theoretical and experimental constraints? If so, how large can they become?

As a concrete example, we consider an aligned 2HDM

Based on

arXiv:2202.03453 (Phys. Rev. Lett.) in collaboration with Henning Bahl and Georg Weiglein

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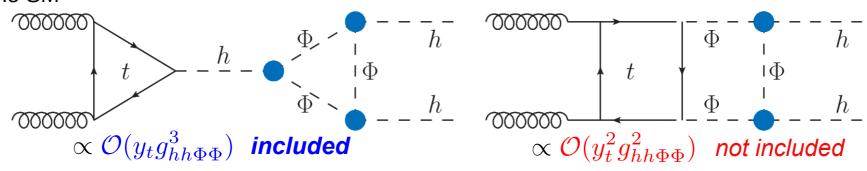
Can we apply di-Higgs results for the aligned 2HDM?

 \succ Current strongest limits on κ_{λ} from ATLAS di-Higgs searches

$$-1.2 < \kappa_{\lambda} < 7.2$$
 [ATLAS-CONF-2024-006]

[where $\kappa_{\lambda} \equiv \lambda_{hhh} / (\lambda_{hhh}^{(0)})^{SM}$]

- What are the assumptions for the ATLAS limits?
 - All other Higgs couplings (to fermions, gauge bosons) are SM-like
 - → this is **ensured by the alignment** ✓
 - The modification of λ_{hhh} is the only source of deviation of the *non-resonant Higgs-pair production cross section* from the SM



- \rightarrow We correctly include all leading BSM effects to di-Higgs production, in powers of $g_{hh\phi\phi}$, up to NNLO! \checkmark
- We can apply the ATLAS limits to our setting!

A parameter scan in the aligned 2HDM

[Bahl, JB, Weiglein PRL '22]

- Our strategy:
 - 1. **Scan BSM parameter space**, keeping only points passing various theoretical and experimental constraints (see below)
 - 2. Identify regions with large BSM deviations in λ_{hhh}
 - 3. Devise a **benchmark scenario** allowing large deviations and investigate impact of experimental limit on λ_{hhh}
- Here: we consider an aligned 2HDM of type-I, but similar results expected for other 2HDM types, or other BSM models with extended Higgs sectors
- Constraints in our parameter scan:
 - 125-GeV Higgs measurements with HiggsSignals
 - Direct searches for BSM scalars with HiggsBounds
 - b-physics constraints, using results from [Gfitter group 1803.01853]

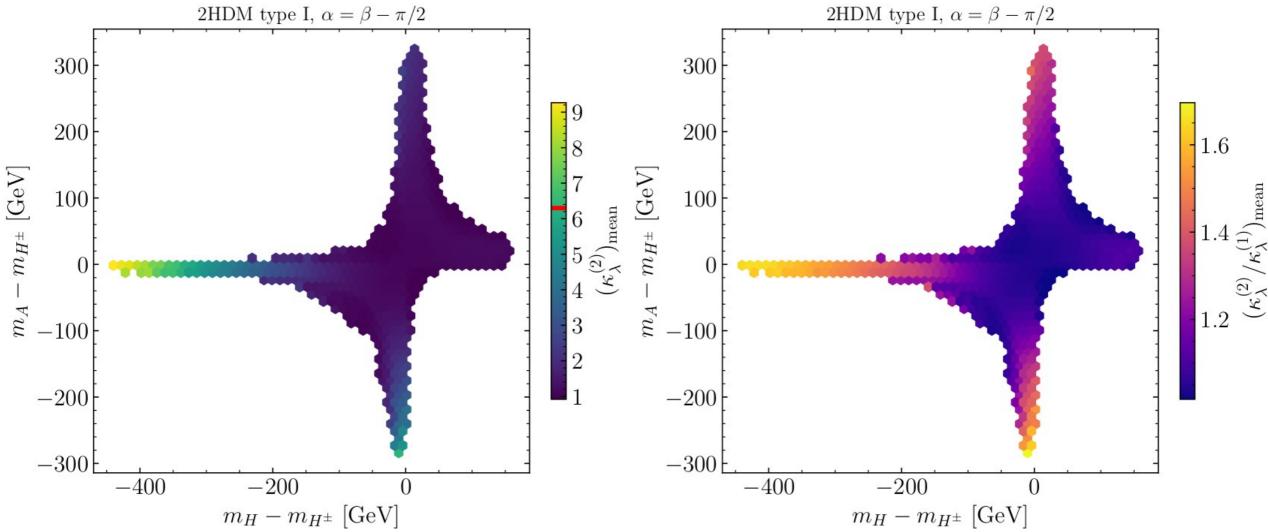
Checked with ScannerS [Mühlleitner et al. 2007.02985]

Checked with ScannerS

- EW precision observables, computed at two loops with THDM_EWPOS [Hessenberger, Hollik '16, '22]
- Vacuum stability
- Boundedness-from-below of the potential
- NLO perturbative unitarity, using results from [Grinstein et al. 1512.04567], [Cacchio et al. 1609.01290]
- For points passing these constraints, we compute κ_{λ} at 1L and 2L, using results from [JB, Kanemura '19]

Parameter scan results

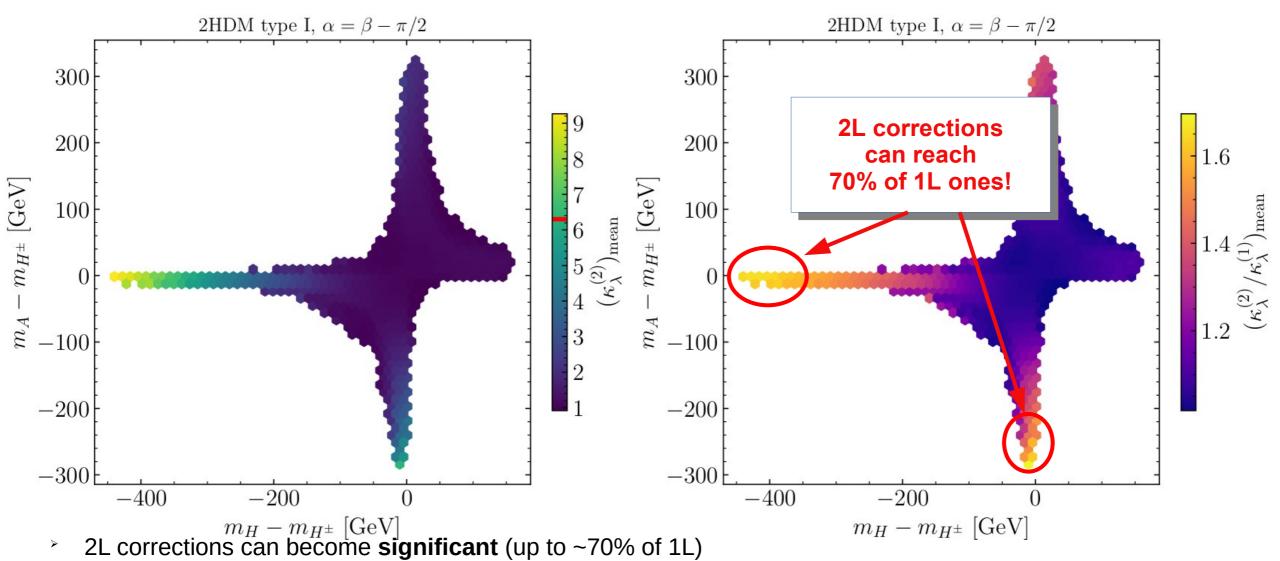
 $\underline{\text{Mean value}} \text{ for } \kappa_{\lambda}^{(2)} = (\lambda_{\text{hhh}}^{(2)})^{\text{2HDM}} / (\lambda_{\text{hhh}}^{(0)})^{\text{SM}} \text{ [left] and } \kappa_{\lambda}^{(2)} / \kappa_{\lambda}^{(1)} = (\lambda_{\text{hhh}}^{(2)})^{\text{2HDM}} / (\lambda_{\text{hhh}}^{(1)})^{\text{2HDM}} \text{ [right] in } (m_{\text{H}} - m_{\text{H}\pm}, m_{\text{A}} - m_{\text{H}\pm}) \text{ plane}$



NB: all previously mentioned constraints are fulfilled by the points shown here

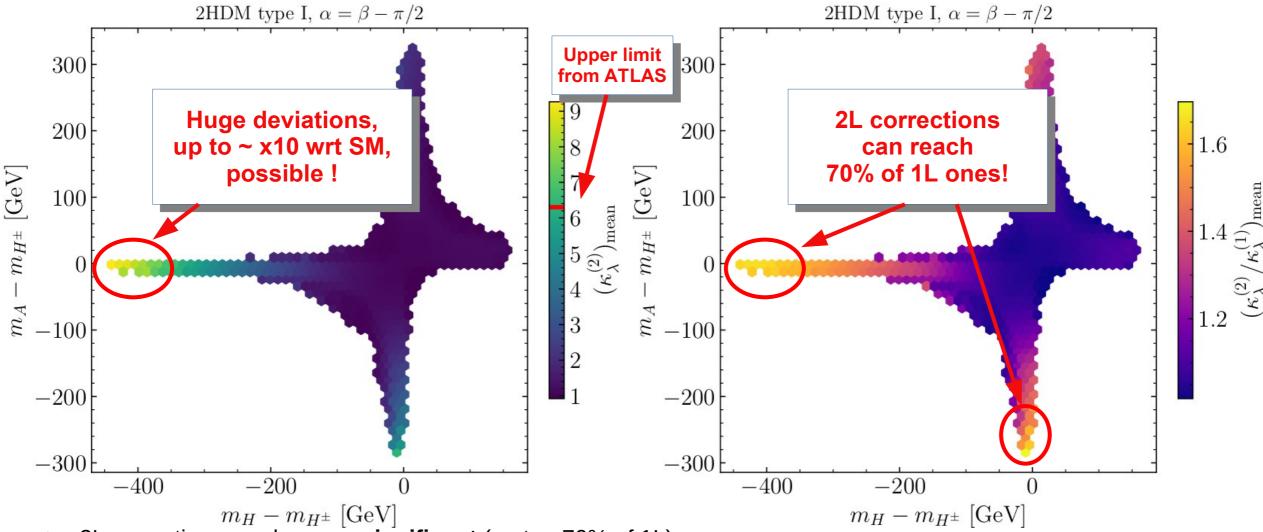
Parameter scan results

 $\underline{\text{Mean value}} \text{ for } \kappa_{\lambda}^{(2)} = (\lambda_{\text{hhh}}^{(2)})^{\text{2HDM}} / (\lambda_{\text{hhh}}^{(0)})^{\text{SM}} \text{ [left] and } \kappa_{\lambda}^{(2)} / \kappa_{\lambda}^{(1)} = (\lambda_{\text{hhh}}^{(2)})^{\text{2HDM}} / (\lambda_{\text{hhh}}^{(1)})^{\text{2HDM}} \text{ [right] in } (m_{\text{H}} - m_{\text{H}\pm}, m_{\text{A}} - m_{\text{H}\pm}) \text{ plane}$



Parameter scan results

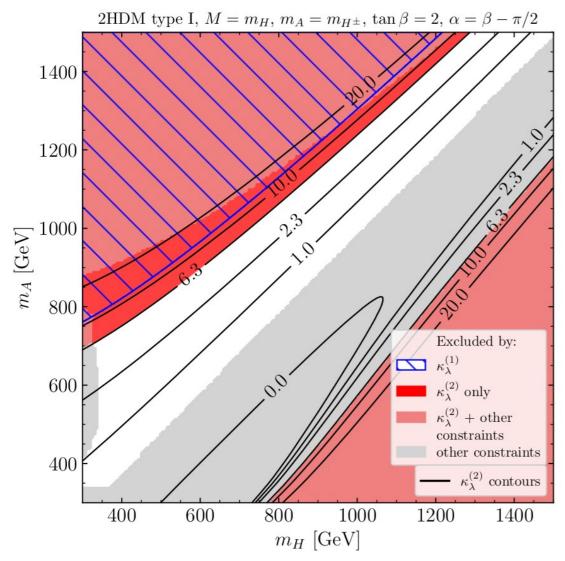
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- 2L corrections can become significant (up to ~70% of 1L)
- Huge enhancements (by a factor ~10) of λ_{hhh} possible for $m_A \sim m_{H\pm}$ and $m_H \sim M$

A benchmark scenario in the aligned 2HDM

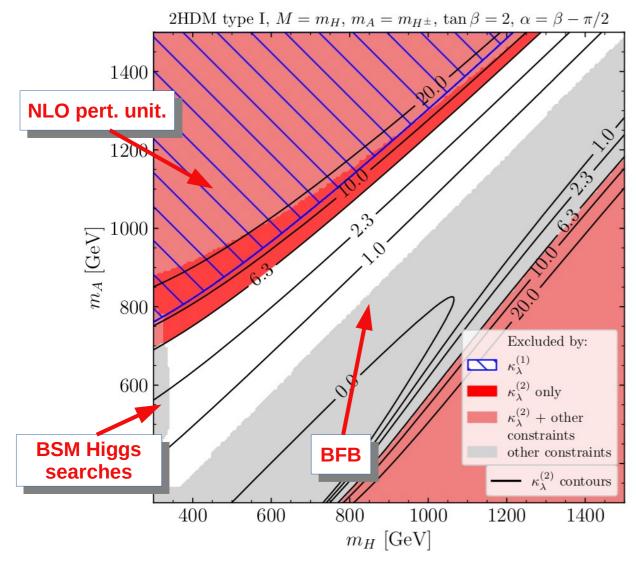
Results shown for aligned 2HDM of type-I, similar for other types (available in backup) We take $m_{A}=m_{H+}$, $M=m_{H}$, $tan\beta=2$



- Grey area: area excluded by other constraints, in particular BSM Higgs searches, boundedness-from-below (BFB), perturbative unitarity
- Light red area: area excluded both by other constraints (BFB, perturbative unitarity) and by $\kappa_{\lambda}^{(2)} > 6.3$ [in region where $\kappa_{\lambda}^{(2)} < -0.4$ the calculation isn't reliable]
- Dark red area: new area that is excluded ONLY by $\kappa_{\lambda}^{(2)} > 6.3$. Would otherwise not be excluded!
- P Blue hatches: area excluded by $κ_λ^{(1)} > 6.3$ → impact of including 2L corrections is significant!

A benchmark scenario in the aligned 2HDM

Results shown for aligned 2HDM of type-I, similar for other types (available in backup) We take $m_{A}=m_{H+}$, $M=m_{H}$, $tan\beta=2$

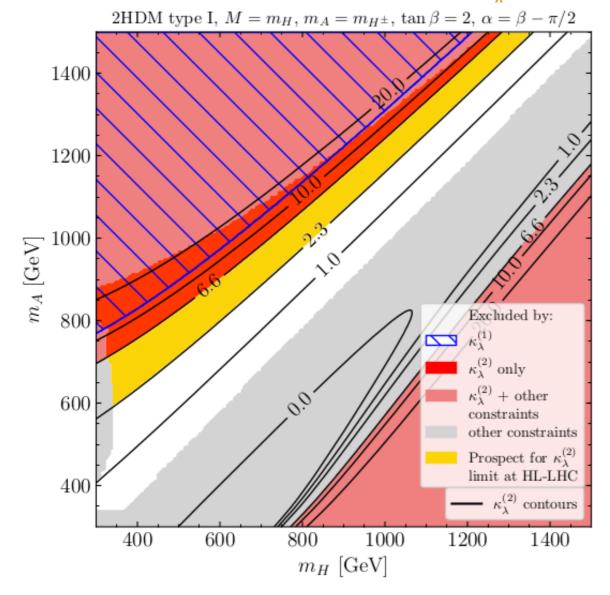


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- Blue hatches: area excluded by $\kappa_{\lambda}^{(1)} > 6.3$ → impact of including 2L corrections is significant!

A benchmark scenario in the aligned 2HDM – future prospects

Suppose for instance the upper bound on κ_{λ} becomes $\kappa_{\lambda} < 2.3$

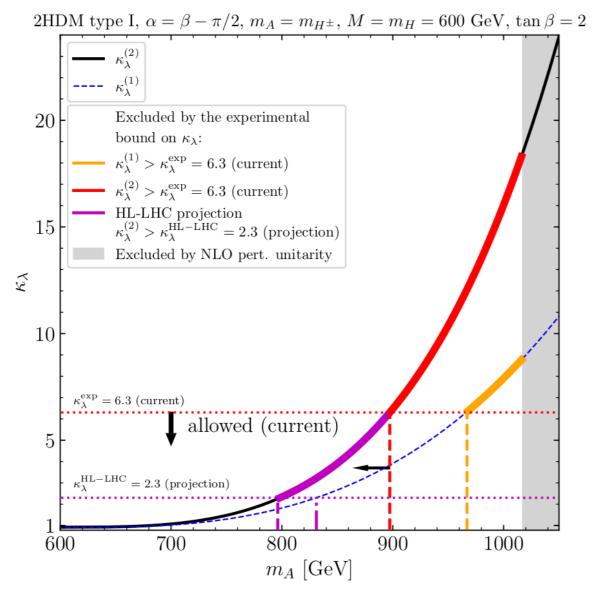
[Bahl, JB, Weiglein '23]



- Fig. 6. Golden area: additional exclusion if the limit on κ_{λ} becomes $\kappa_{\lambda}^{(2)} < 2.3$ (achievable at HL-LHC)
- Of course, prospects even better with an e+ecollider!
- Experimental constraints, such as Higgs physics, may also become more stringent, however **not** theoretical constraints (like BFB or perturbative unitarity)

A benchmark scenario in the aligned 2HDM - 1D scan

Within the previously shown plane, we fix $M=m_{H}=600$ GeV, and vary $m_{A}=m_{H\pm}$

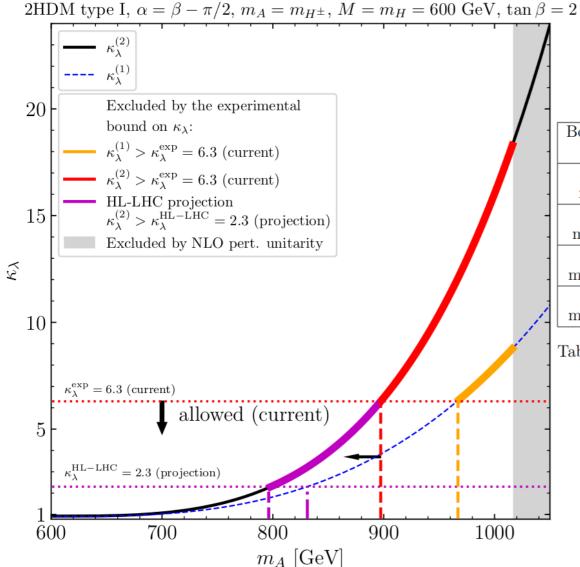


[Bahl, JB, Weiglein PRL '22]

- Illustrates the significantly improved reach of the experimental limit when including **2L corrections** in calculation of κ_{λ}
- A stricter choice for the perturbative unitarity constraint (grey) does not significantly change the region excluded by $\kappa_{\lambda}^{(2)}$

A benchmark scenario in the aligned 2HDM – 1D scan



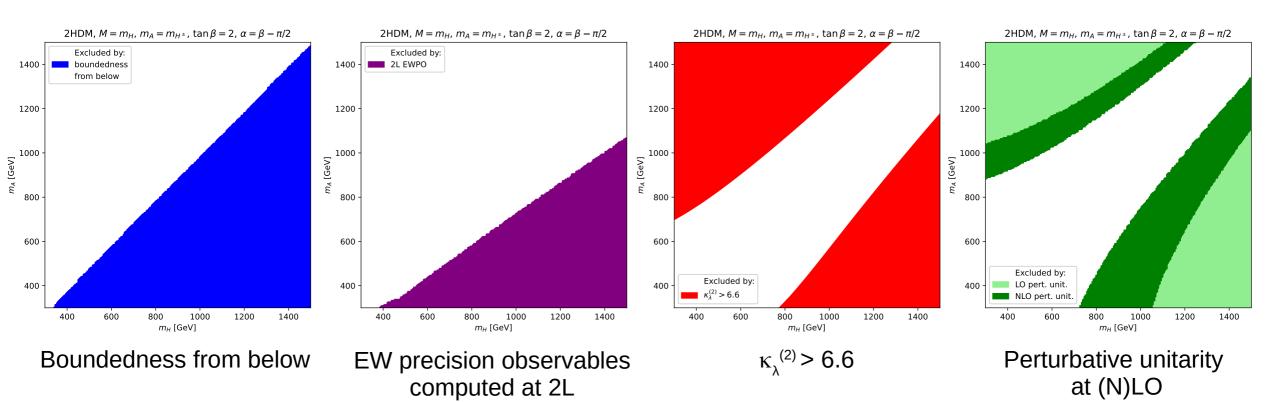


Bound on eigenvalues	$\max(m_A)$ with	$\max(m_A)$ with	$\max(m_A)$ with
	LO pert. unit.	NLO pert. unit.	with finite $\sqrt{s} \in [3 \text{ TeV}, 10 \text{ TeV}]$
$\max(a_i) < 1$	1161 GeV	$1017 \; \mathrm{GeV}$	_
$\max(\mathfrak{Re}(a_i)) < 1$	1161 GeV	1033 GeV	1260 GeV
$\max(a_i) < 0.5$	917 GeV	937 GeV	_
$\max(\mathfrak{Re}(a_i)) < 0.5$	917 GeV	958 GeV	929 GeV
$\max(a_i) < 0.49$	911 GeV	933 GeV	_
$\max(\mathfrak{Re}(a_i)) < 0.49$	911 GeV	956 GeV	922 GeV
$\max(a_i) < 0.45$	889 GeV	912 GeV	_
$\max(\mathfrak{Re}(a_i)) < 0.45$	889 GeV	948 GeV	897 GeV

Table 1: Maximal values of m_A allowed in the benchmark scenario under the constraint of perturbative unitarity, at LO and NLO, and for different upper bounds on the $2 \to 2$ scattering eigenvalues used in the perturbative unitarity constraint. Note that tree-level scattering eigenvalues are all real, so there is no difference between using max or $\Re \mathfrak{e}(\max)$ for the left column.

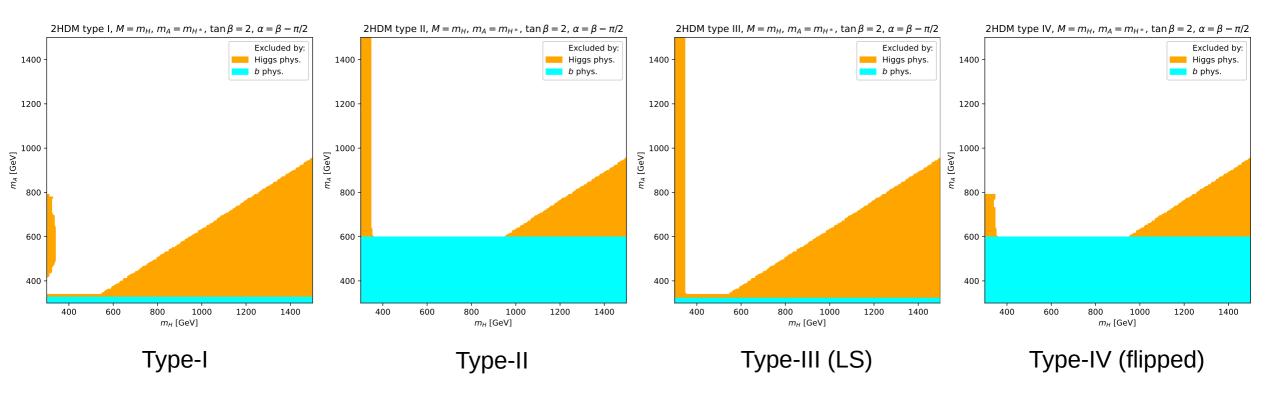
2HDM benchmark plane – individual theoretical constraints

Constraints shown below are independent of 2HDM type



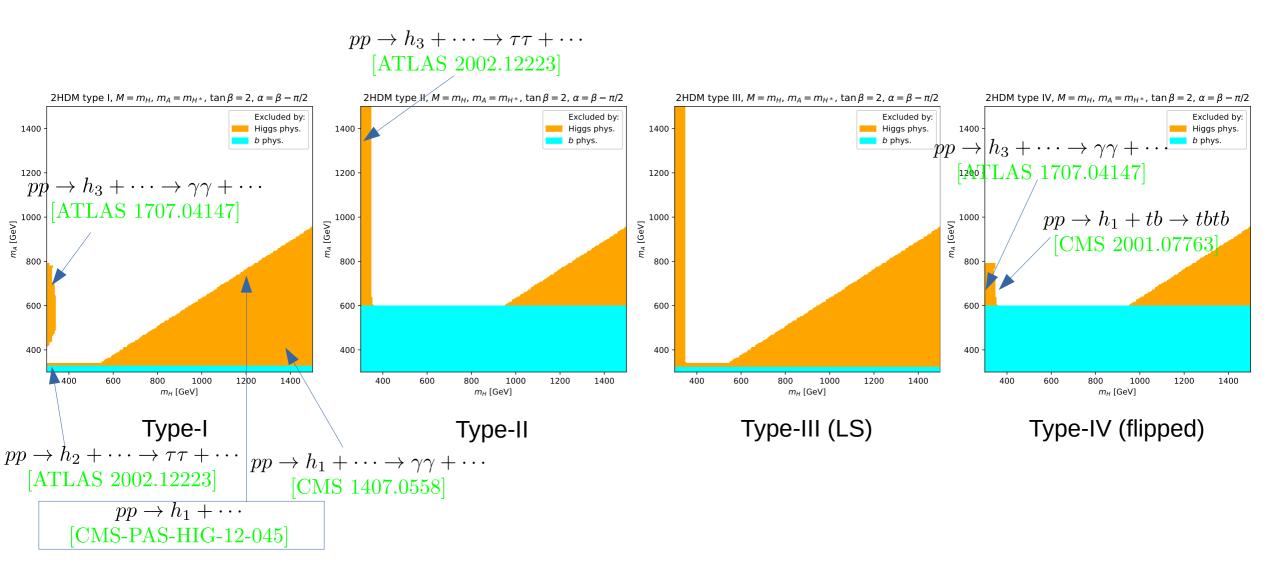
2HDM benchmark plane – experimental constraints

i.e. Higgs physics (via HiggsBounds and HiggsSignals) and b physics (from [Gfitter group 1803.01853])

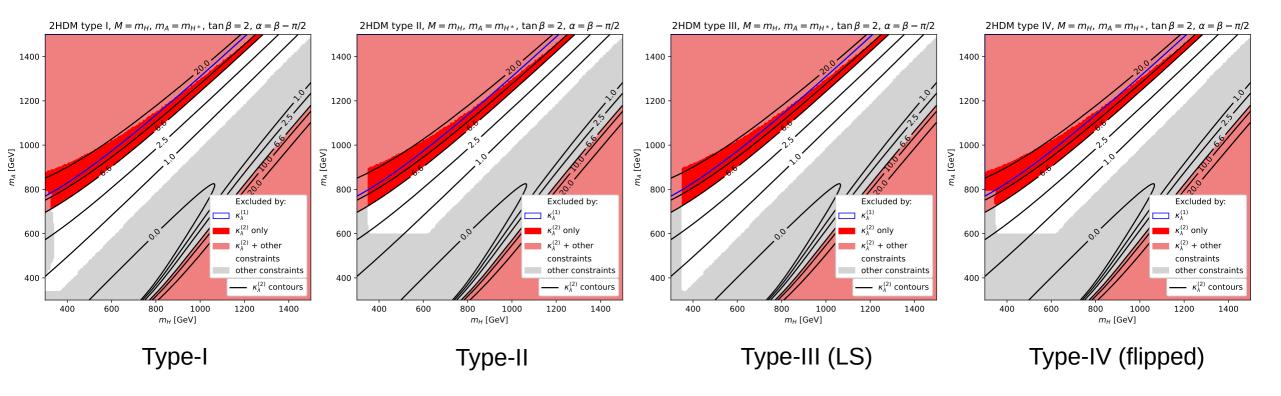


2HDM benchmark plane – experimental constraints

i.e. Higgs physics (via HiggsBounds and HiggsSignals) and b physics (from [Gfitter group 1803.01853])



2HDM benchmark plane – results for all types



anyH3: full 1L calculation of λ_{hhh} in any renormalisable model

